



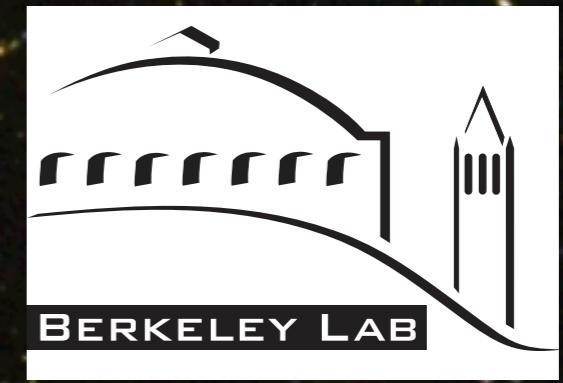
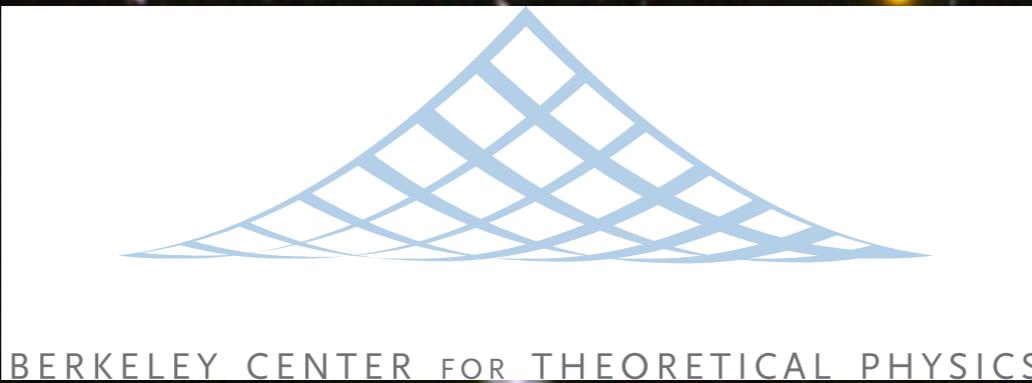
INSTITUTE FOR THE PHYSICS AND  
MATHEMATICS OF THE UNIVERSE



東京大学  
THE UNIVERSITY OF TOKYO

# Inflation and Particle Physics

LiteBIRD Kick-off Symposium @ ISAS, JAXA  
Hitoshi Murayama (Berkeley, Kavli ITPMU) July 1, 2019



BERKELEY CENTER FOR THEORETICAL PHYSICS



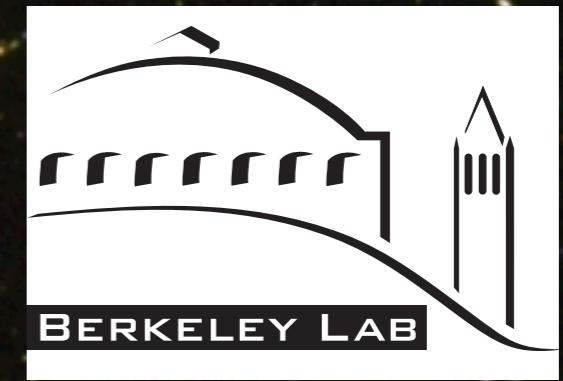
INSTITUTE FOR THE PHYSICS AND  
MATHEMATICS OF THE UNIVERSE



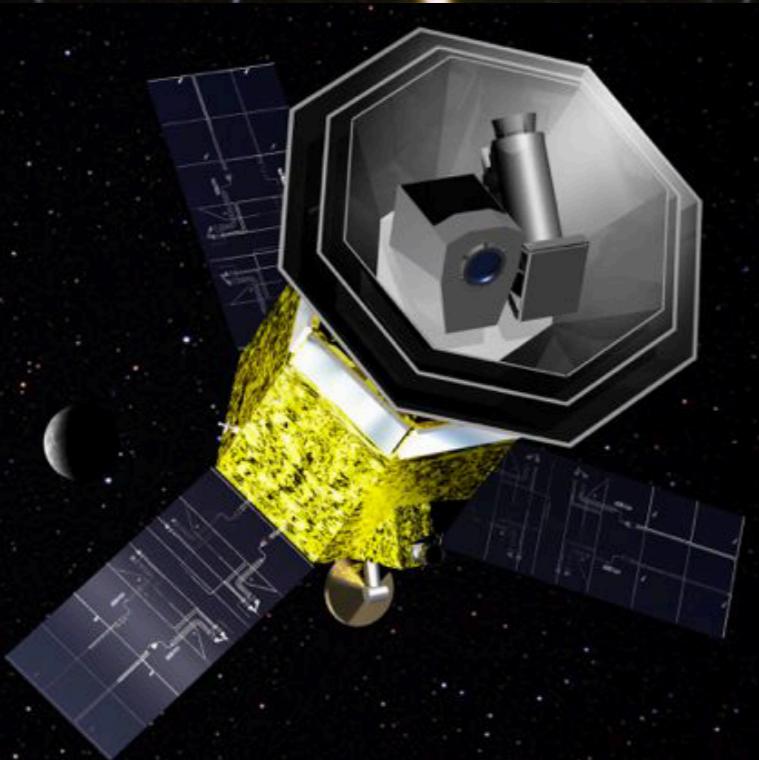
東京大学  
THE UNIVERSITY OF TOKYO

# Inflation for Pedestrians

LiteBIRD Kick-off Symposium @ ISAS, JAXA  
Hitoshi Murayama (Berkeley, Kavli IPMU) July 1, 2019



BERKELEY CENTER FOR THEORETICAL PHYSICS

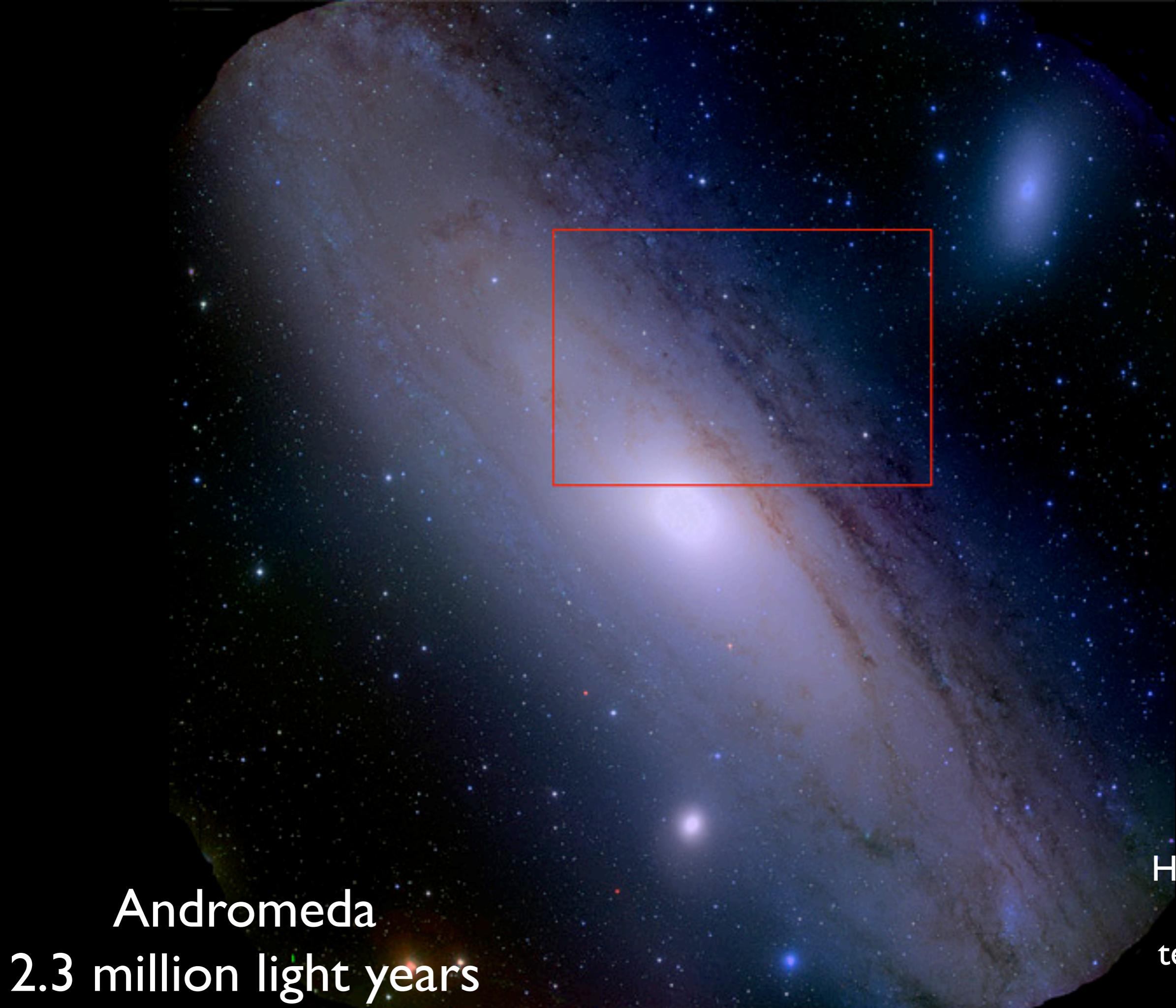


How did the Universe begin?  
What is its fate?  
What is it made of?  
How does it work?  
Where do we come from?



A dark, star-filled background representing space, with numerous small, glowing points of light scattered across the frame.

Telescopes are  
time machines



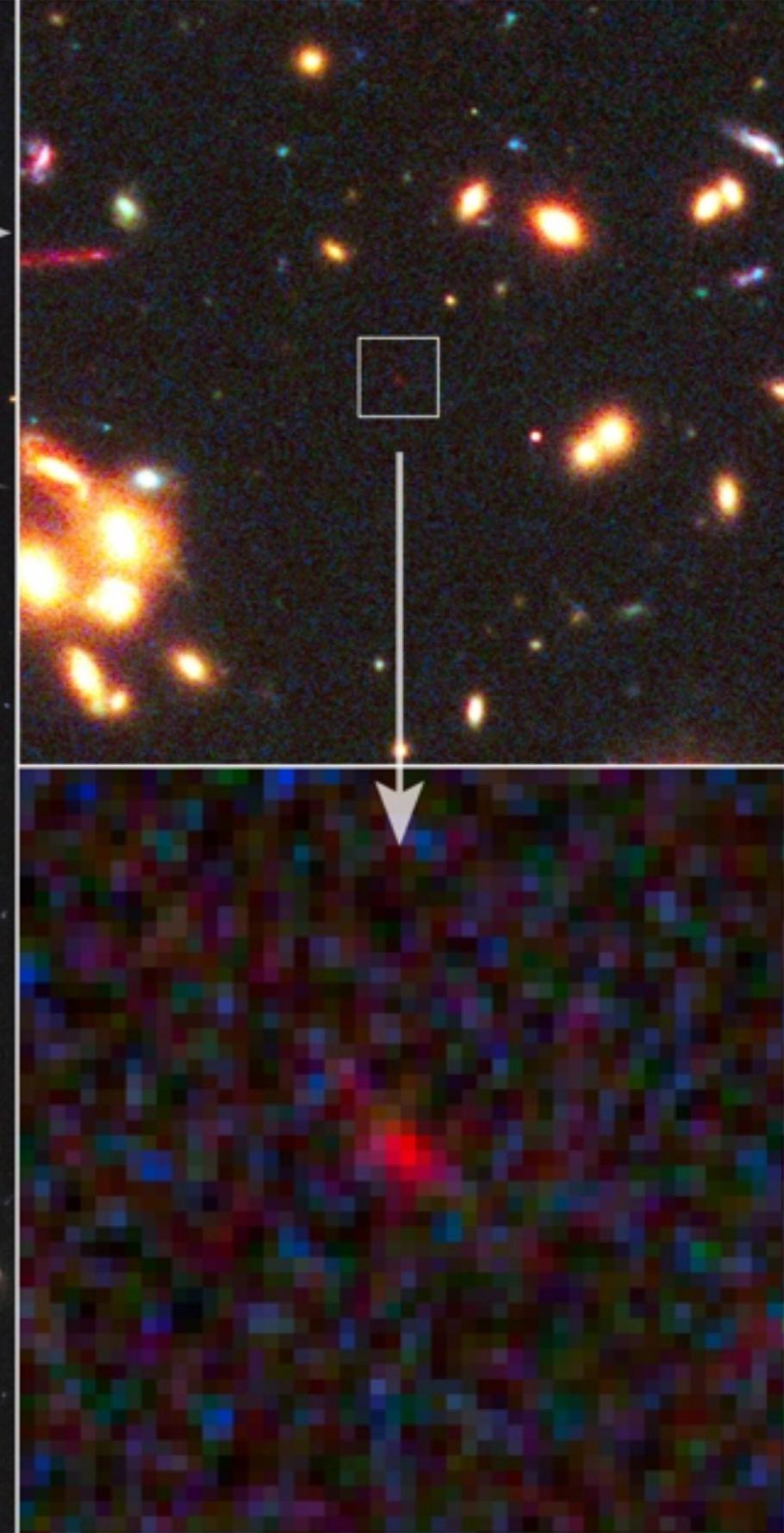
Andromeda  
2.3 million light years

HSC team  
Subaru  
telescope

cluster of galaxies  
2.1 billion light year



Abell 2218



Galaxy Cluster MACS J1149+2223

galaxy @ 13.3 billion light years

A Distant Gravitationally-Lensed Galaxy at Redshift = 9.6

NASA / ESA / STScI / J. Hora (Harvard-Smithsonian CfA)

High-Redshift Galaxy MACS1149-JD

Hubble Space Telescope • ACS • WFC3

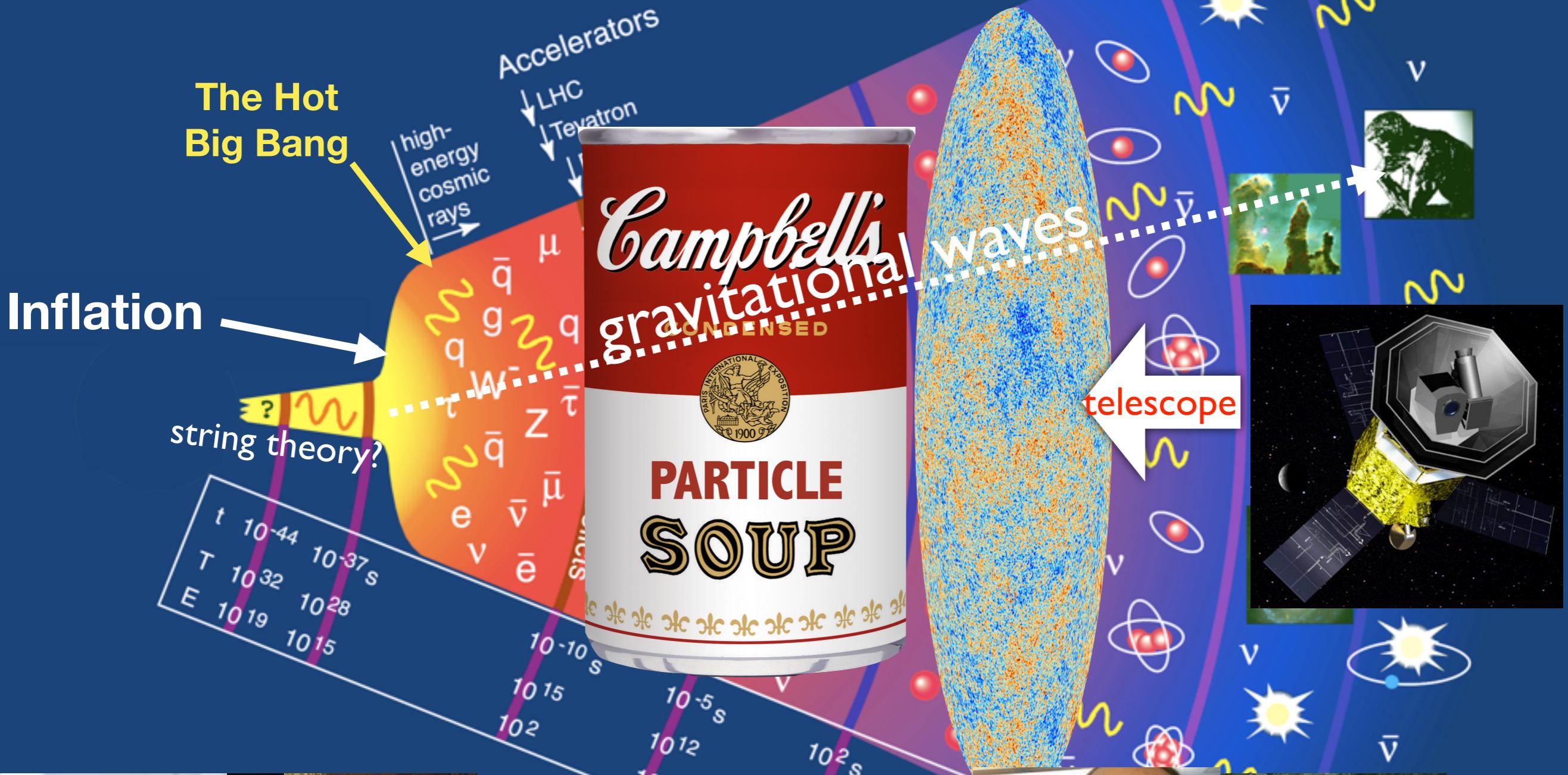
ssc2012-12a

# Picture of the Big Bang

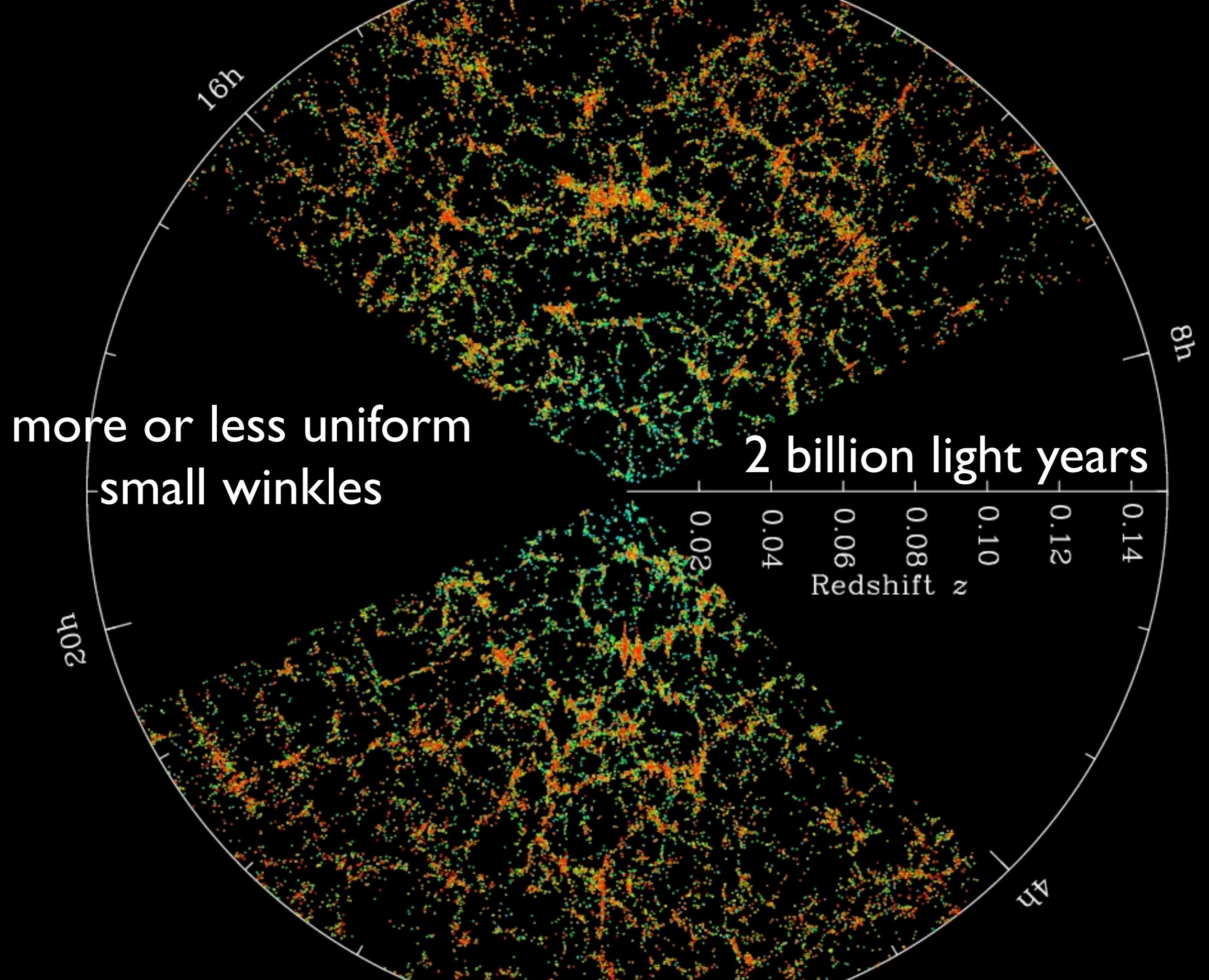
Temperature variation  $\sim 1/100,000$



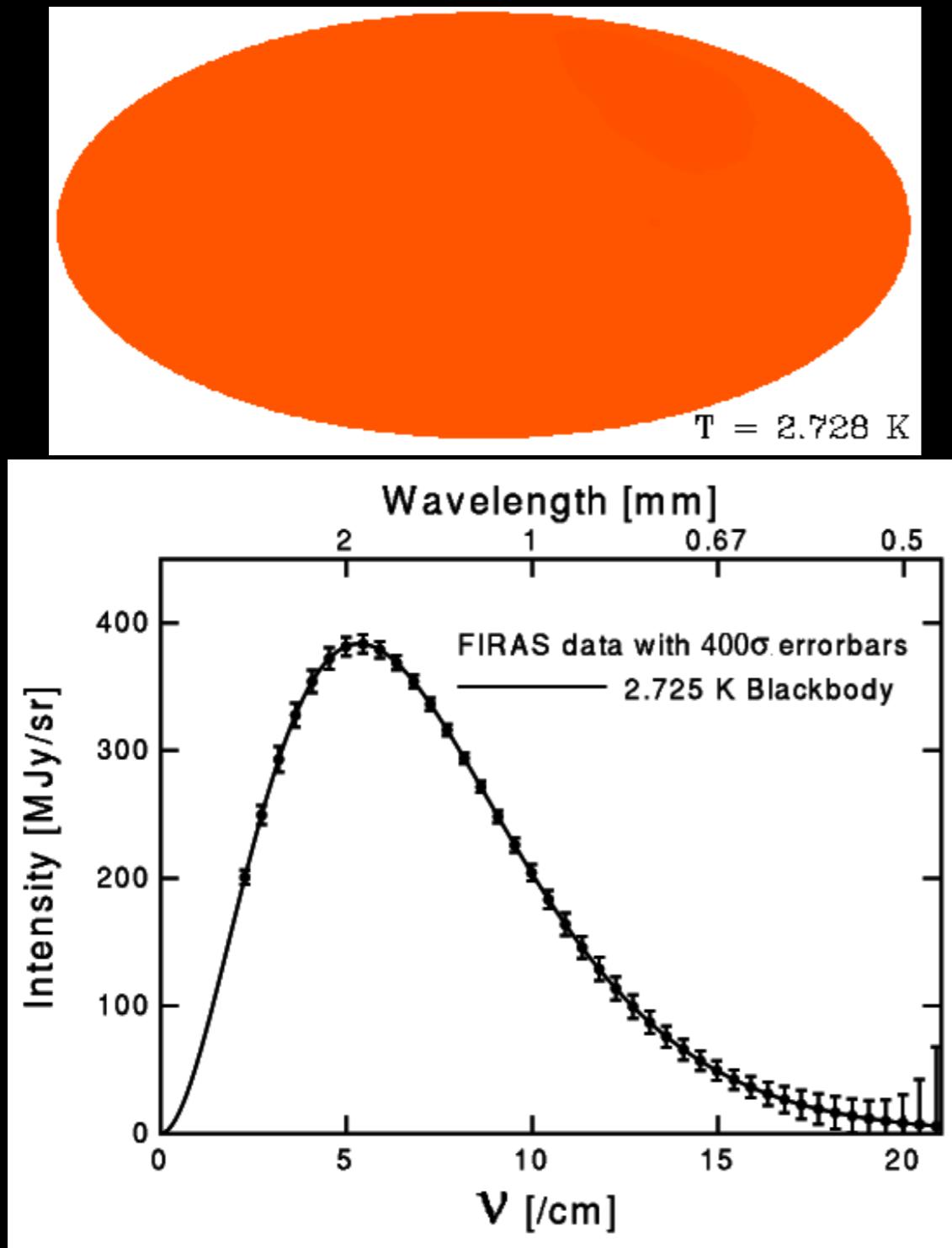
# History of the Universe







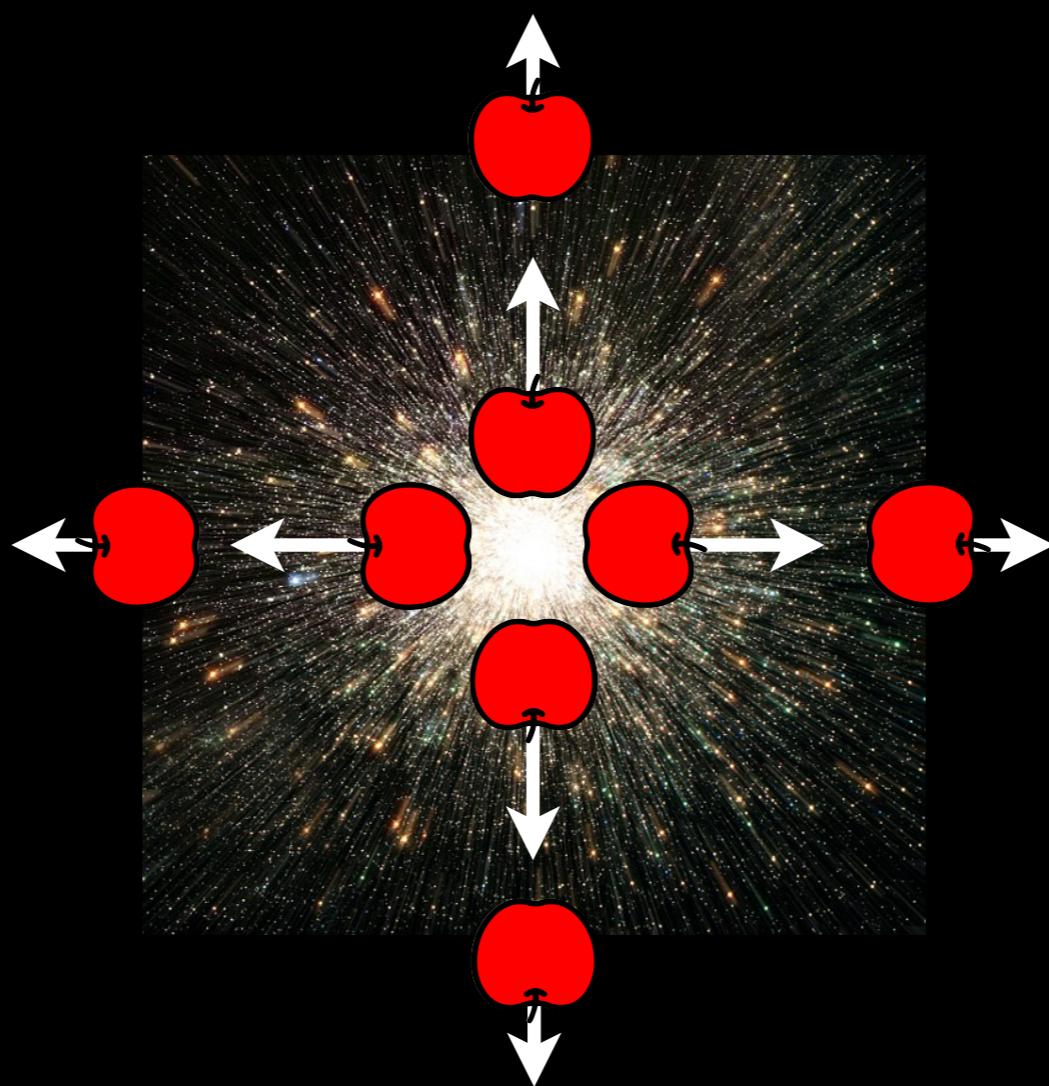
# How do they know each other?



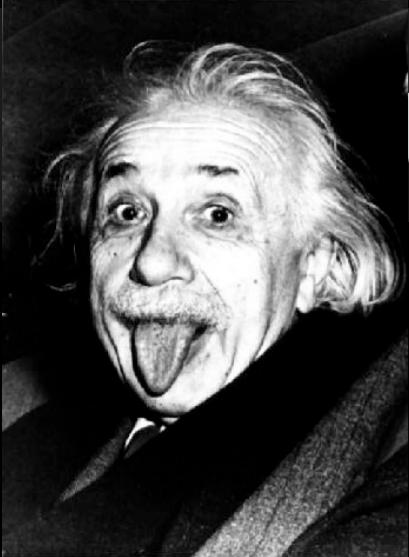
- Like having discovered two remote islands in very different parts of the world, but people speak the same language
- we suspect they were together at some point

**Cosmic Microwave  
Background  
(CMB)**

# Expansion should slow down



Gravity only pulls, never pushes



# How is that possible?

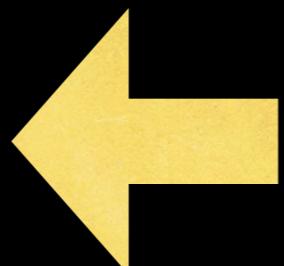
Einstein's equation

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = 8\pi G_N T_{\mu\nu}$$

How Universe expands

what is inside

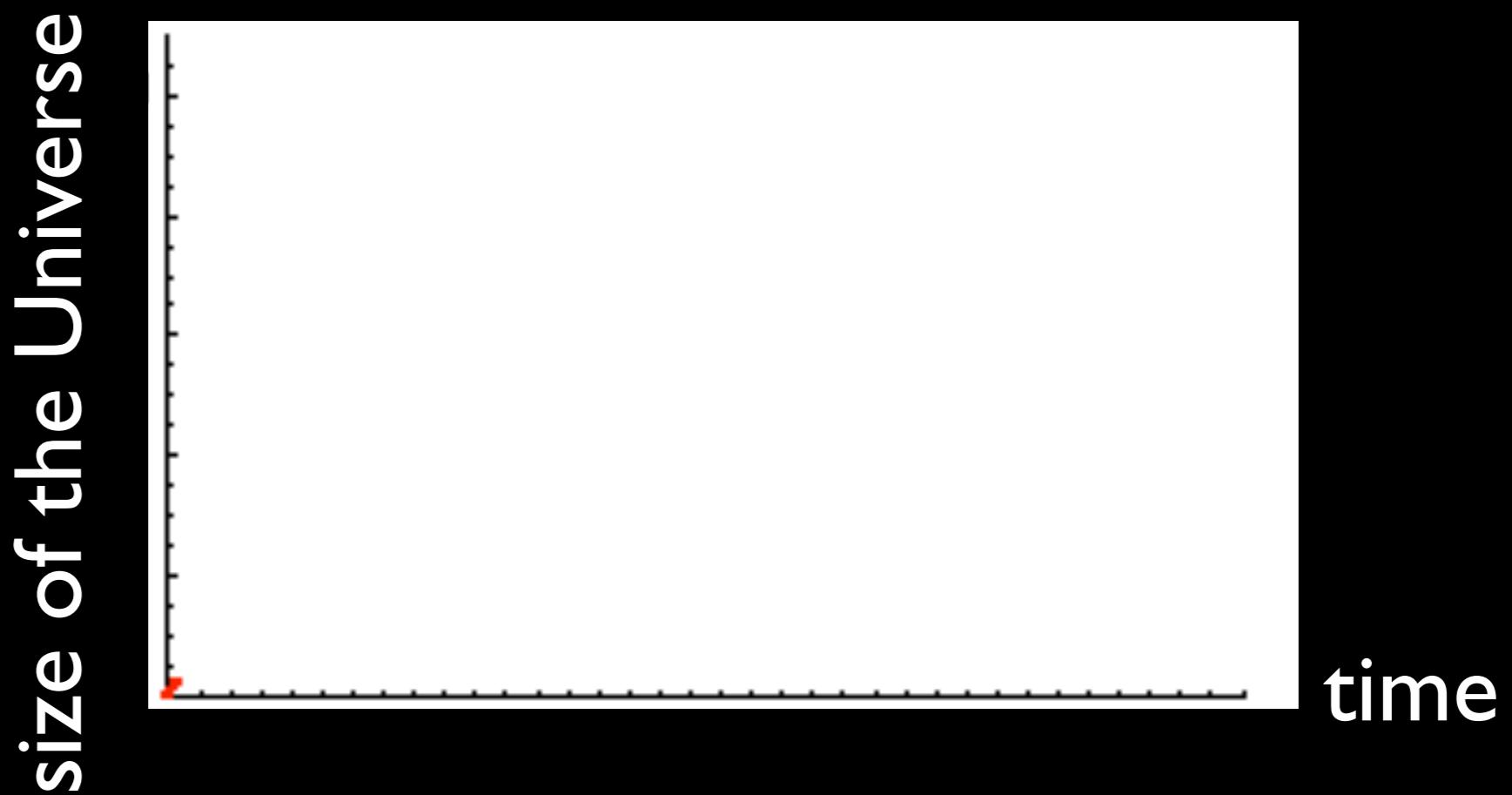
faster expansion



more energy

# How does it happen?

more or less constant energy in a given volume

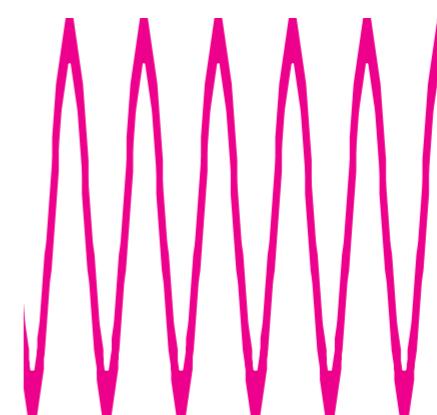


# near sighted

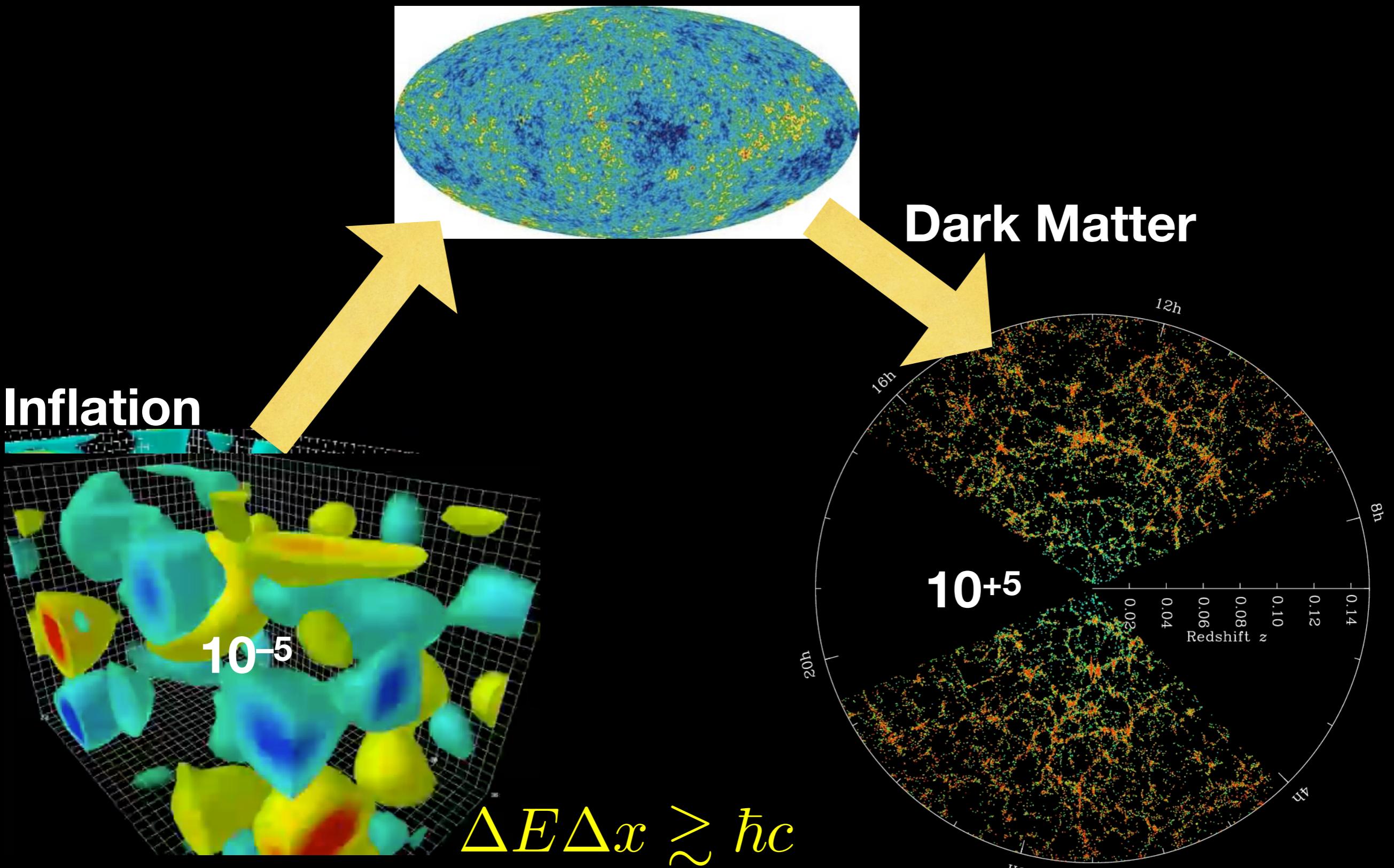
- What you are seeing one moment is *gone* by inflation the next moment
- feel very near-sighted
- “horizon”
- small space

⇒ big uncertainty

$$\Delta E \Delta x \gtrsim \hbar c$$



# Seeds for structure



# Inflation sowed seeds Dark Matter grew them



without dark matter



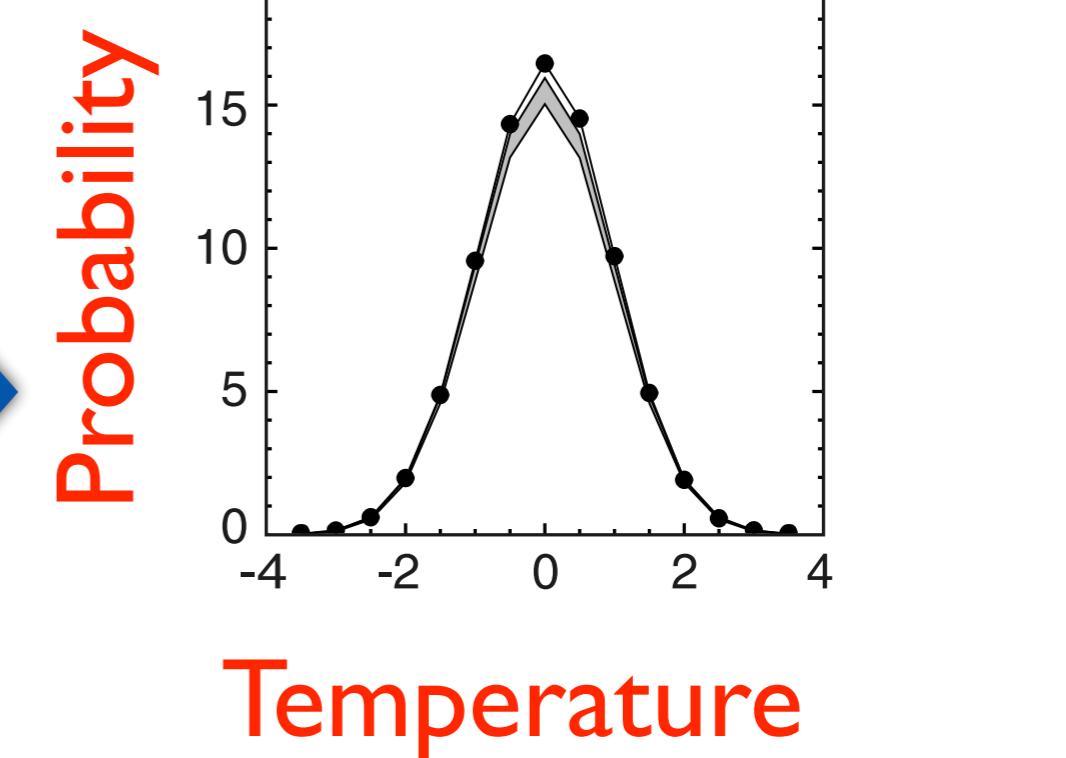
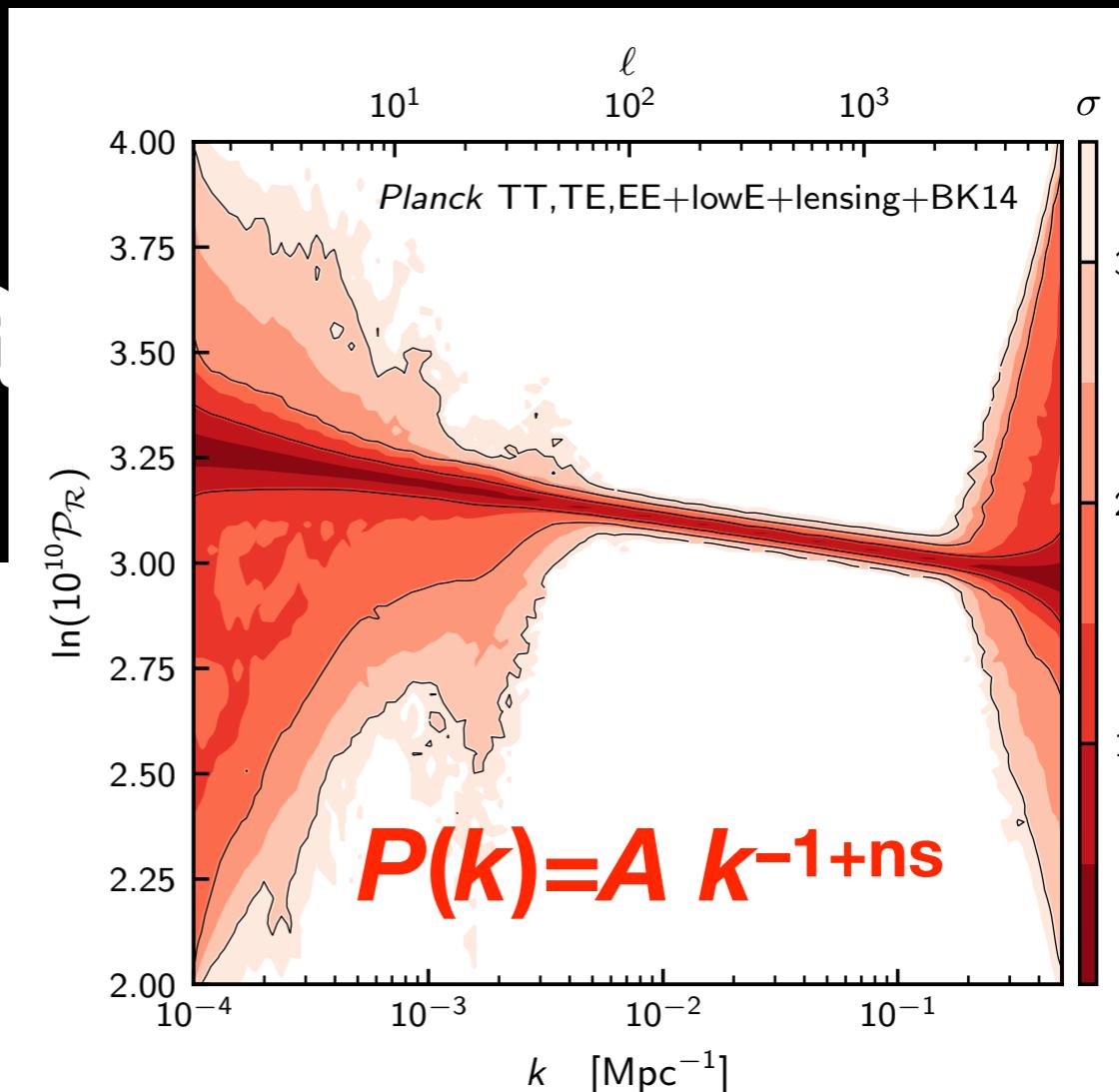
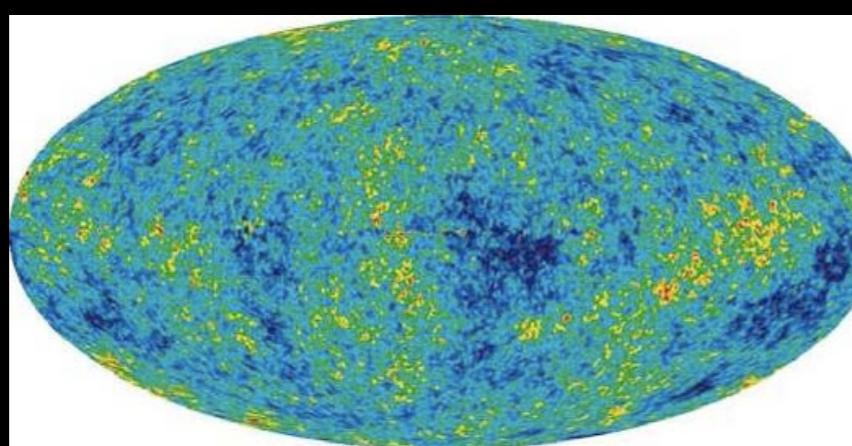
with dark matter



Inflation is Dad  
Dark Matter is Mom

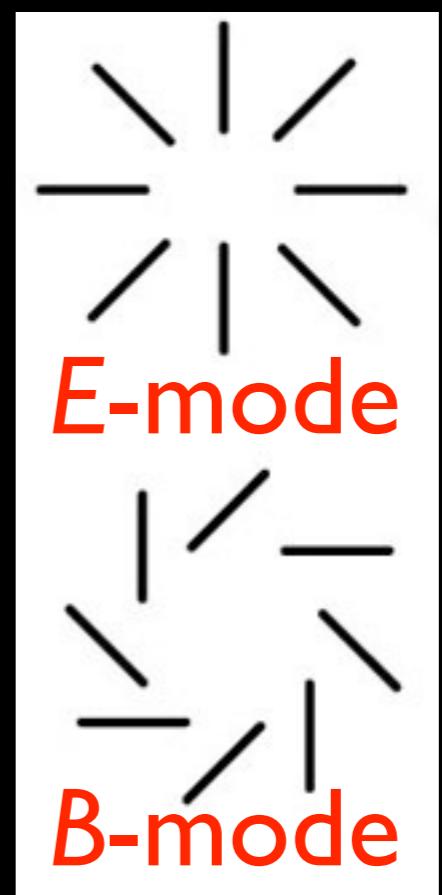
# Looks planckian

- Uncertainty principle leads to Gaussian distribution for the simplest cases
- Indeed!
- Matches simplest models of inflation



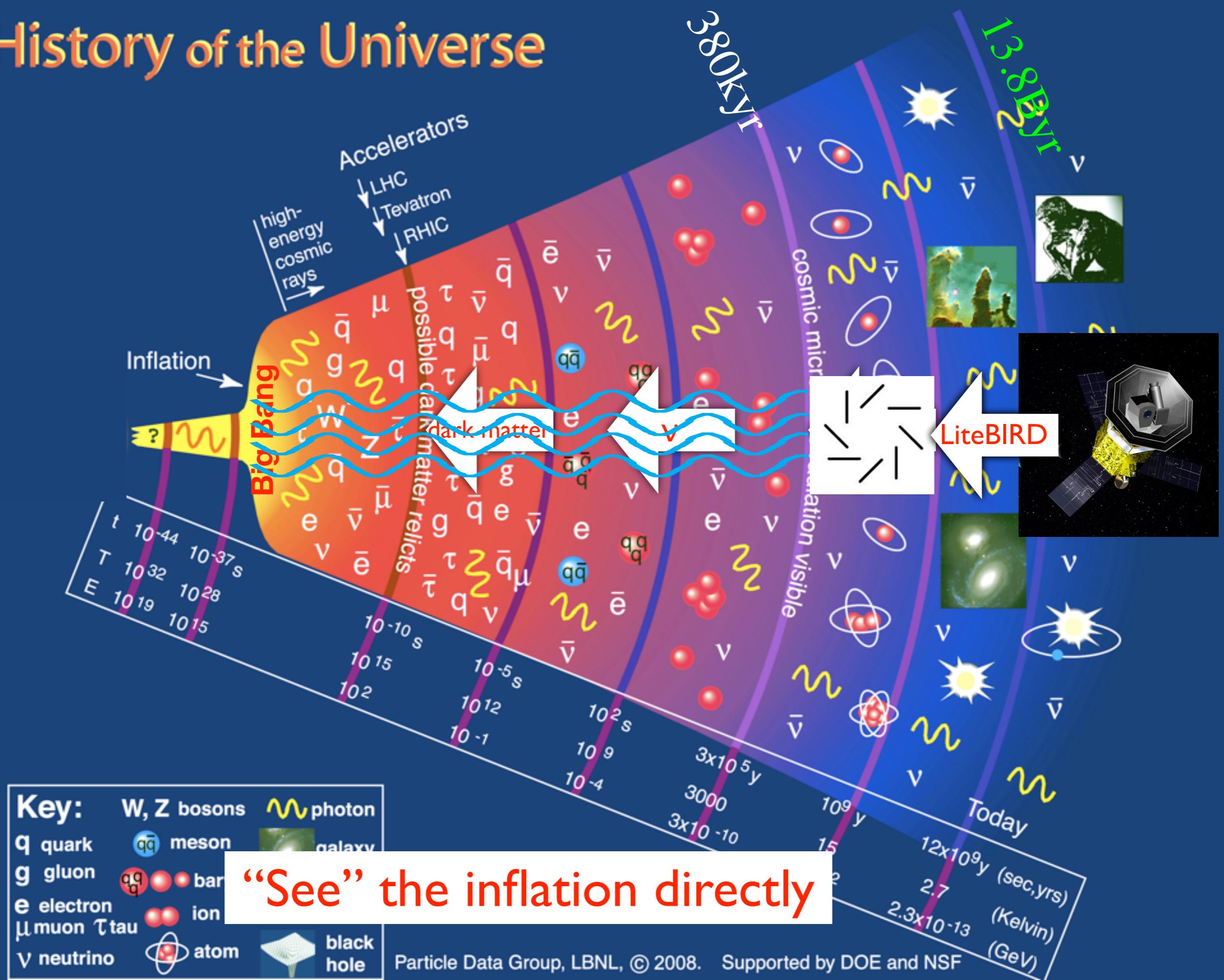
# Still many mysteries

- What caused inflation?
- Which moment?
- How much?
- Definitive proof?



polarization

# History of the Universe

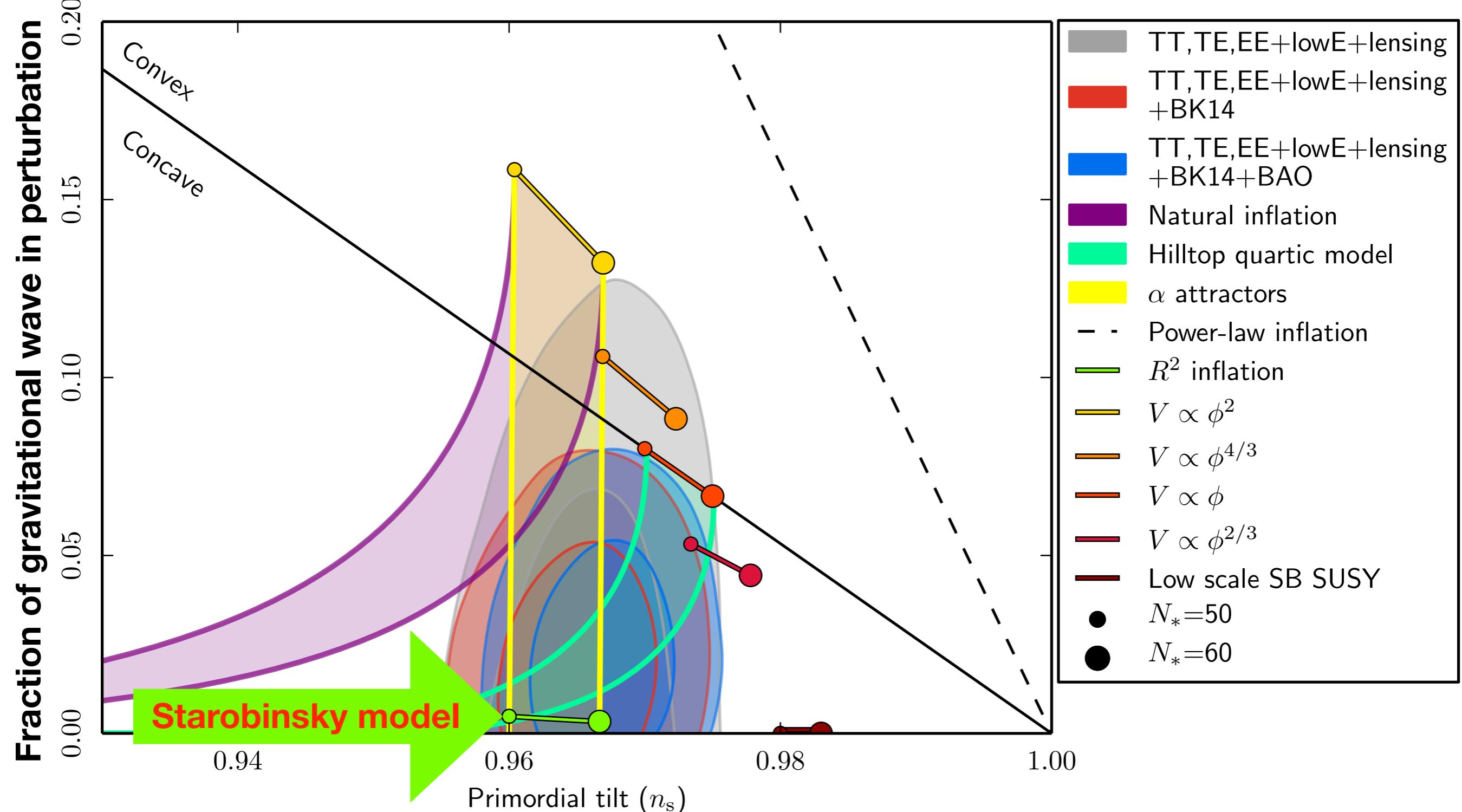


# Gravitational Wave exists!



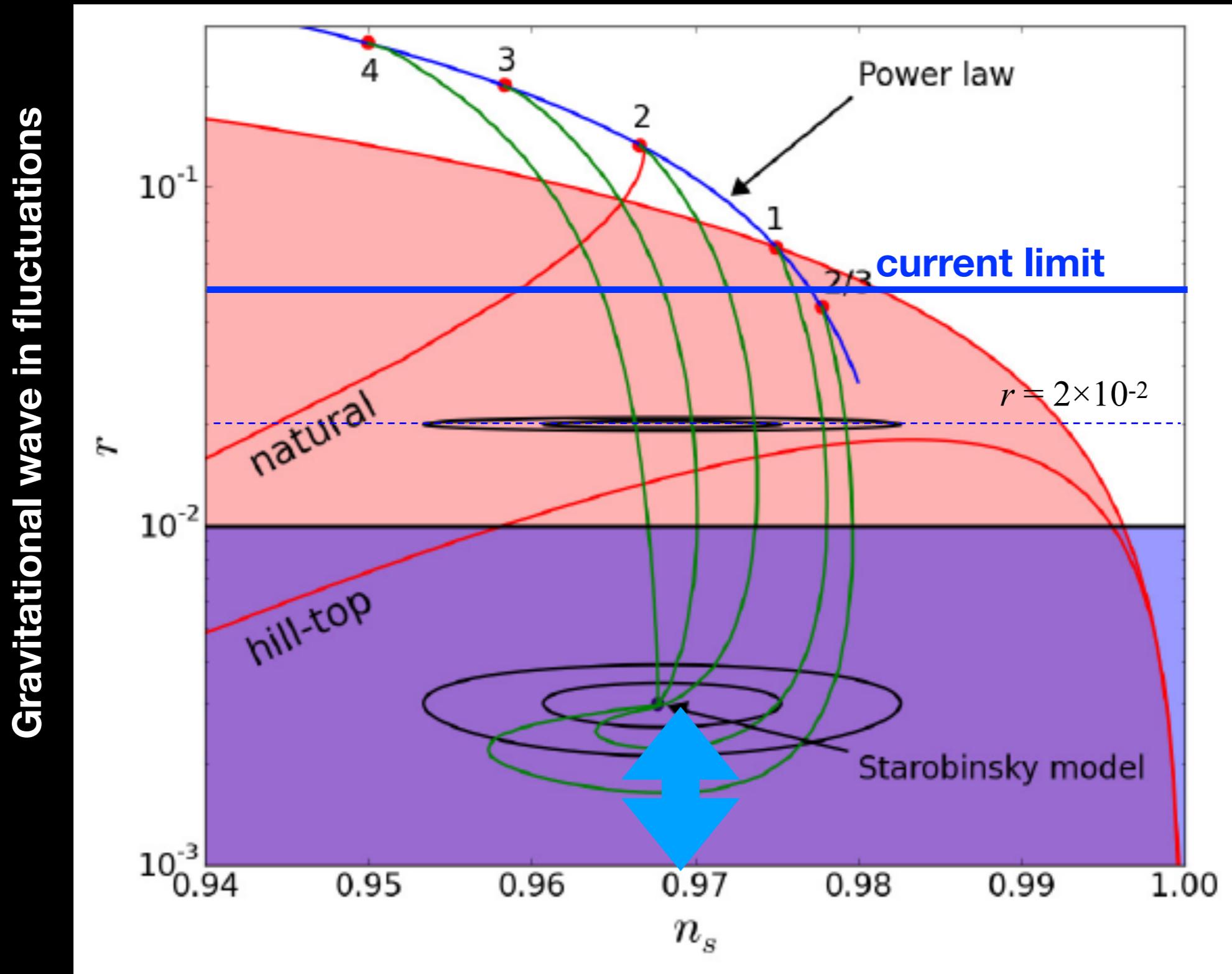
Scale of Effect Vastly Exaggerated

The whole Earth stretches only by  $10^{-14}\text{cm}$



$$P(k)=A k^{-1+n_s}$$

# What we expect with LiteBIRD



We absolutely need  $\delta r = 10^{-3}$

# String Theory

- Inflaton field cannot go over a large distance  $\Delta\phi \lesssim M_{\text{Planck}}$

H. Ooguri and C. Vafa, hep-th/0605264

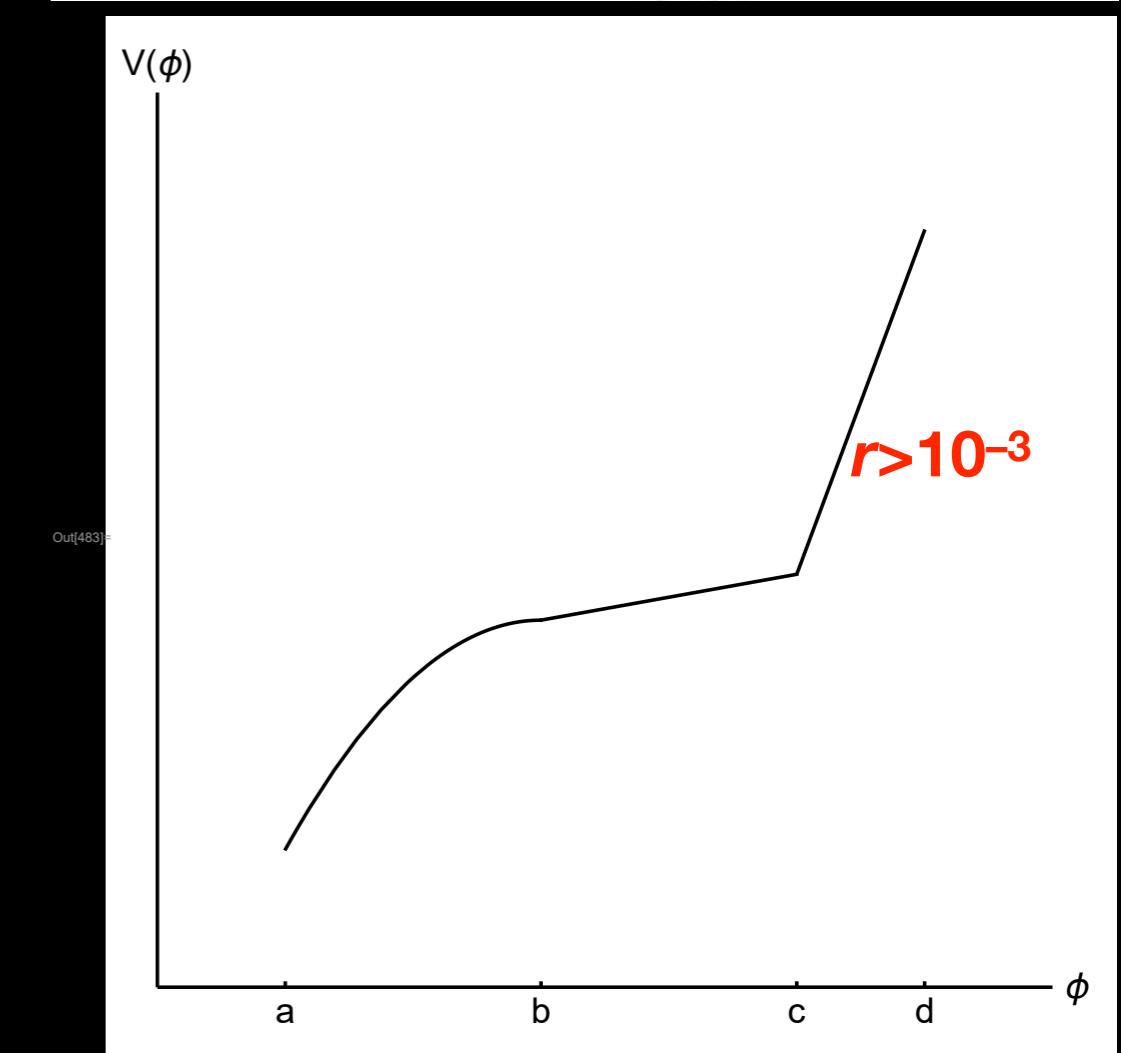
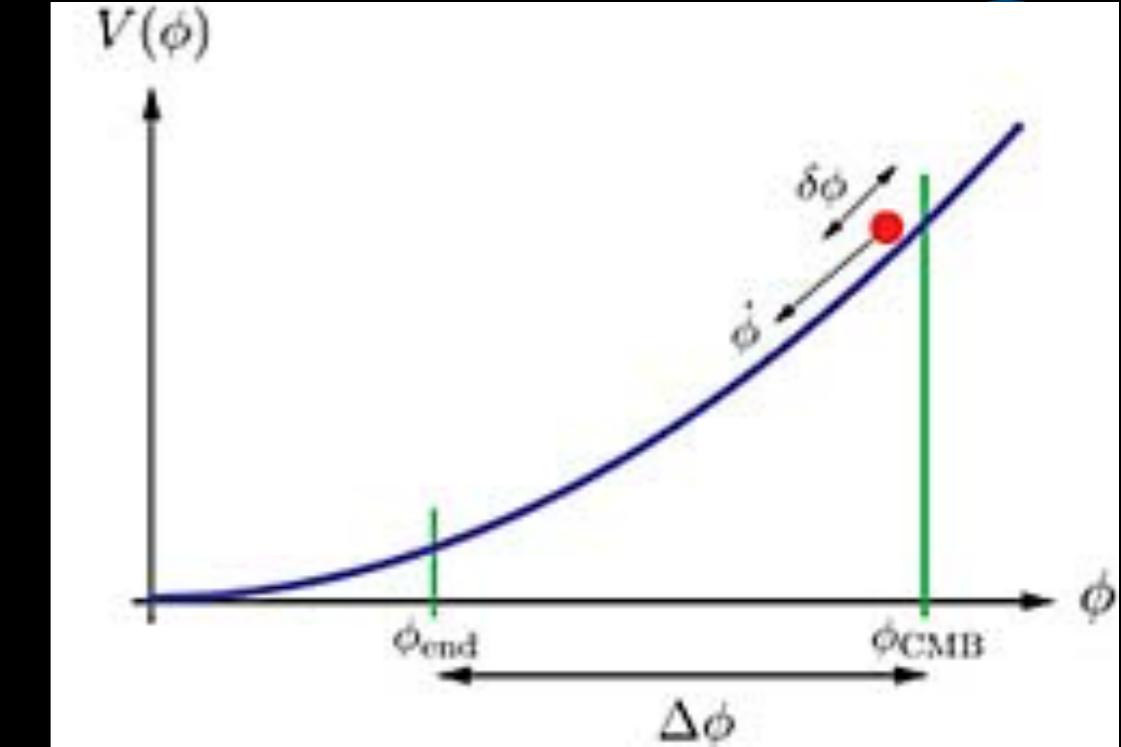
- can't expect large tensor component

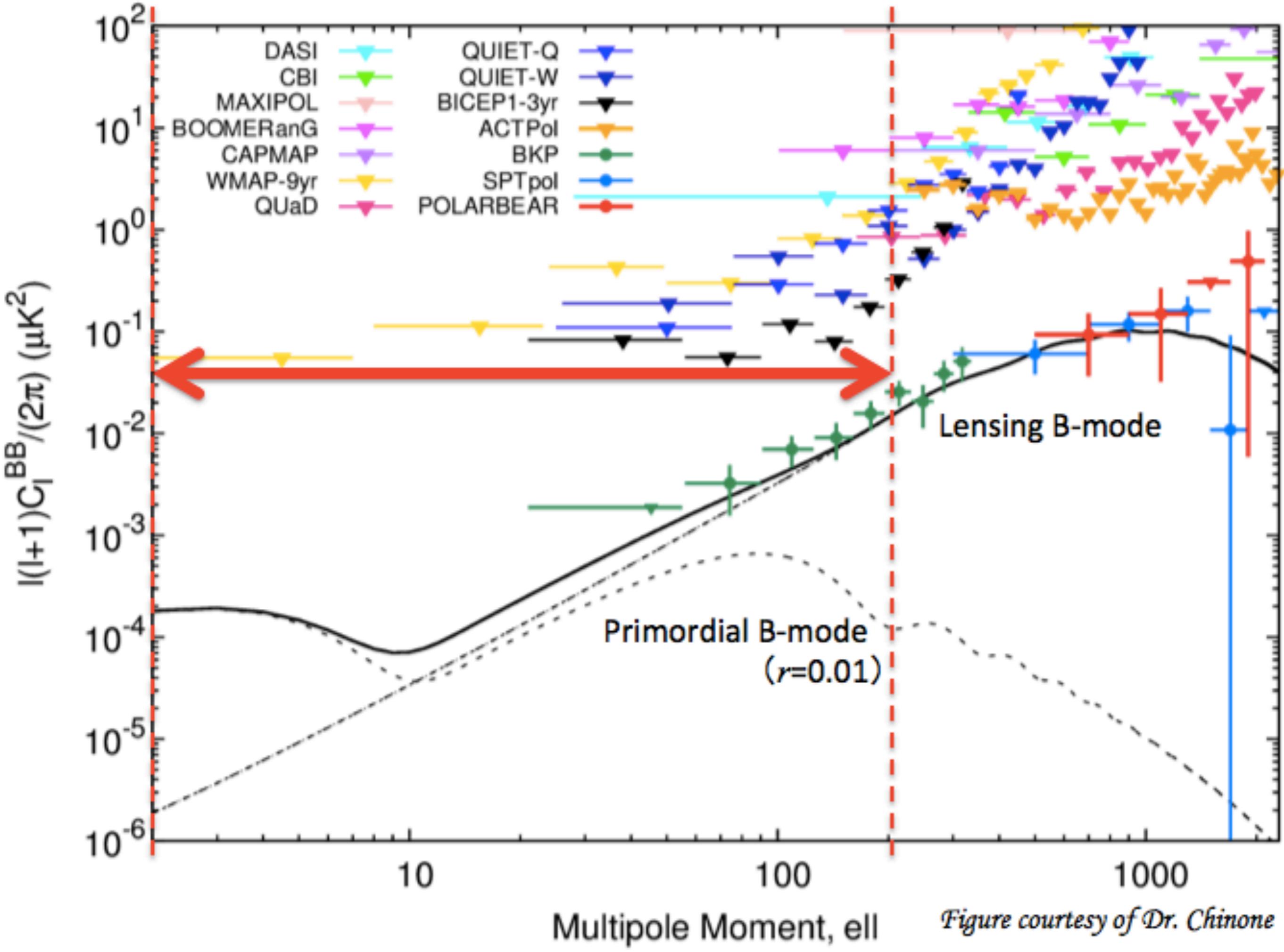
$$r \approx 8 \left( \frac{d\phi}{dN} \right)^2 \lesssim 10^{-3}$$

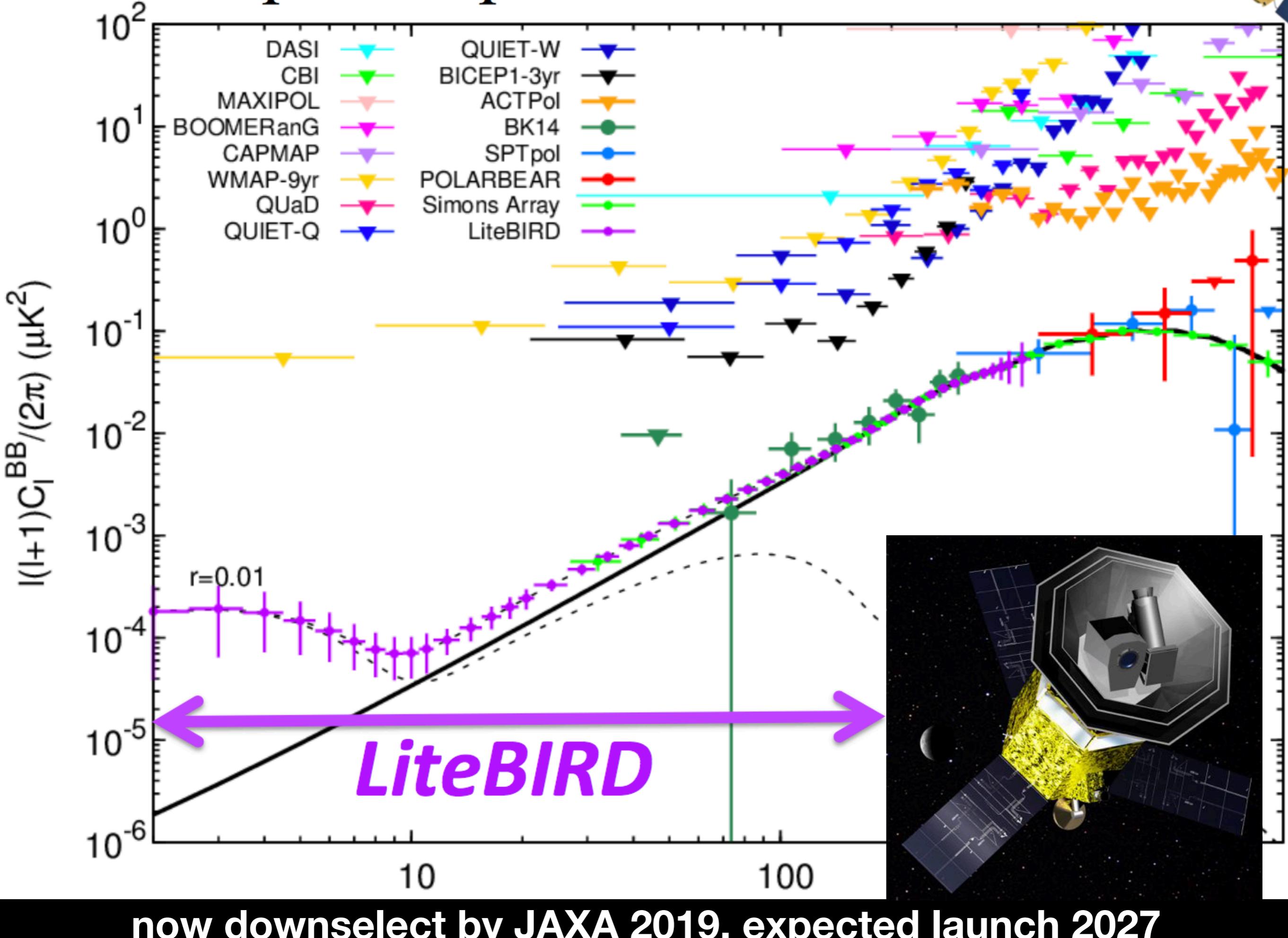
- bad news for CMB S4, LiteBIRD?
- Inflation may move substantially only for several initial e-folds
- Can expect much larger  $r$ !
- Still observation will put string theory on the edge

HM, Masahito Yamazaki and Tsutomu T. Yanagida, arXiv:1809.00478

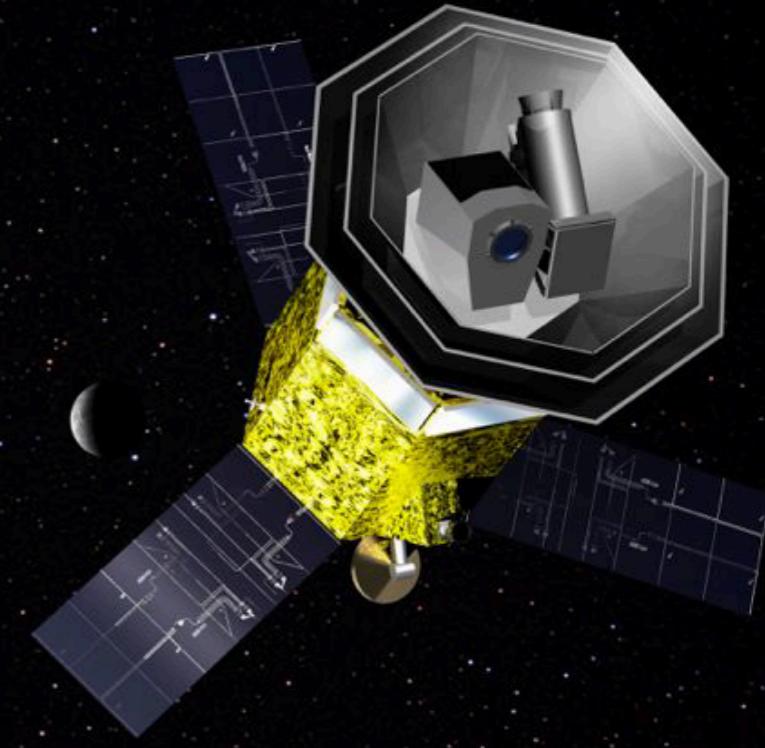
Chien-I Chiang, Jacob Leedom, HM, arXiv:1811.01987



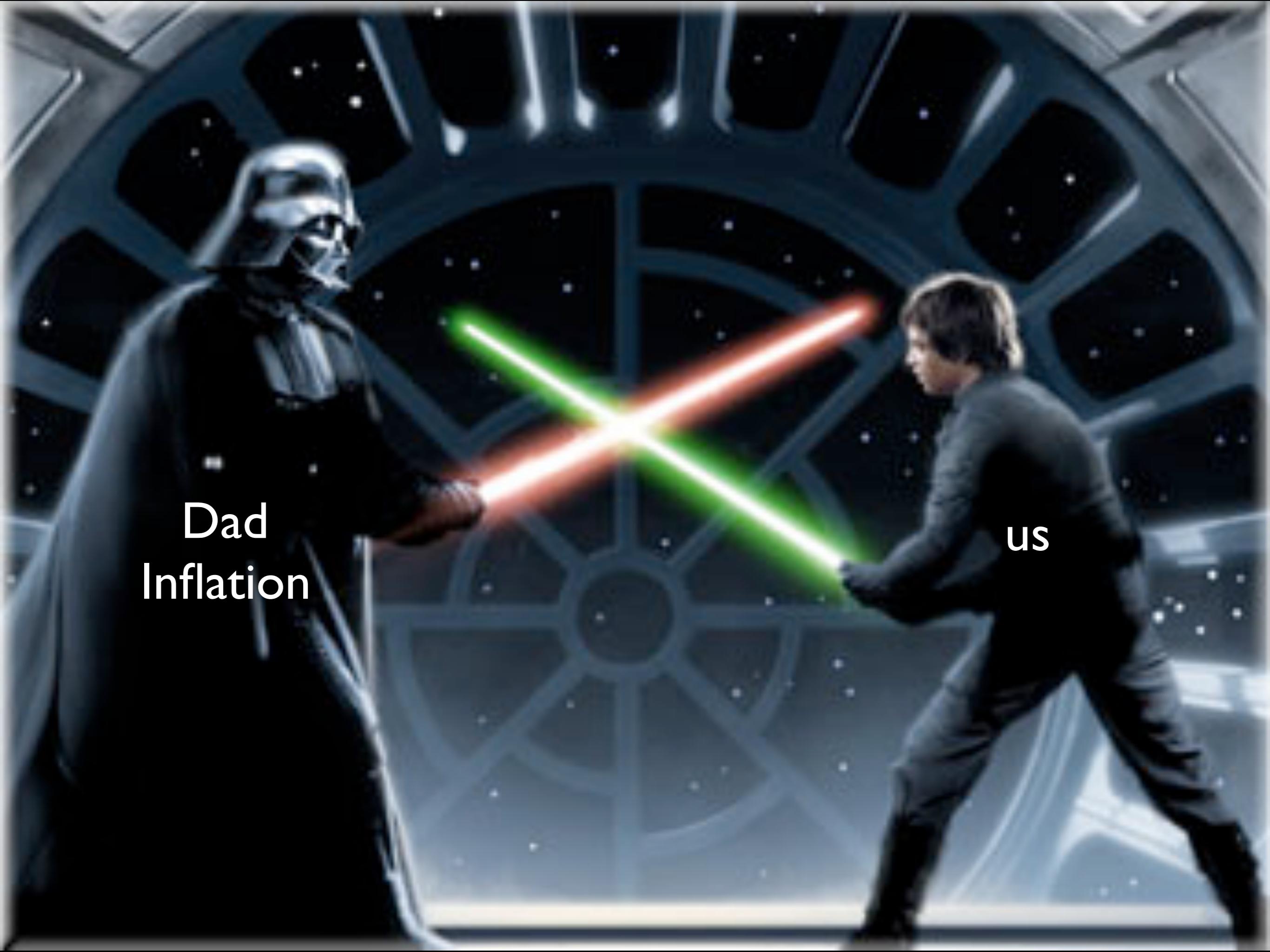




# LiteBIRD



- CMB B-mode polarization is unique  
*Universe before the Big Bang*
- We can prove inflation (and meet our Dad!)
- specify “who” was our Dad
- probe theories that unify gravity and quantum physics



Dad  
Inflation

us