

2016/10/31

7th KIAS workshop on Cosmology and Structure Formation

LiteBIRD

Lite satellite for the studies of B-mode polarization
and Inflation from cosmic background Radiation Detection

Ryo Nagata
with LiteBIRD collaboration

Probing cosmic inflation with cosmic microwave background (CMB) polarization

Primordial gravitational waves

Primordial CMB B-mode
curl patterns as “fingerprints” of inflation

Beginning of
our universe

Inflation

Hot
Big Bang

Recombi-
nation

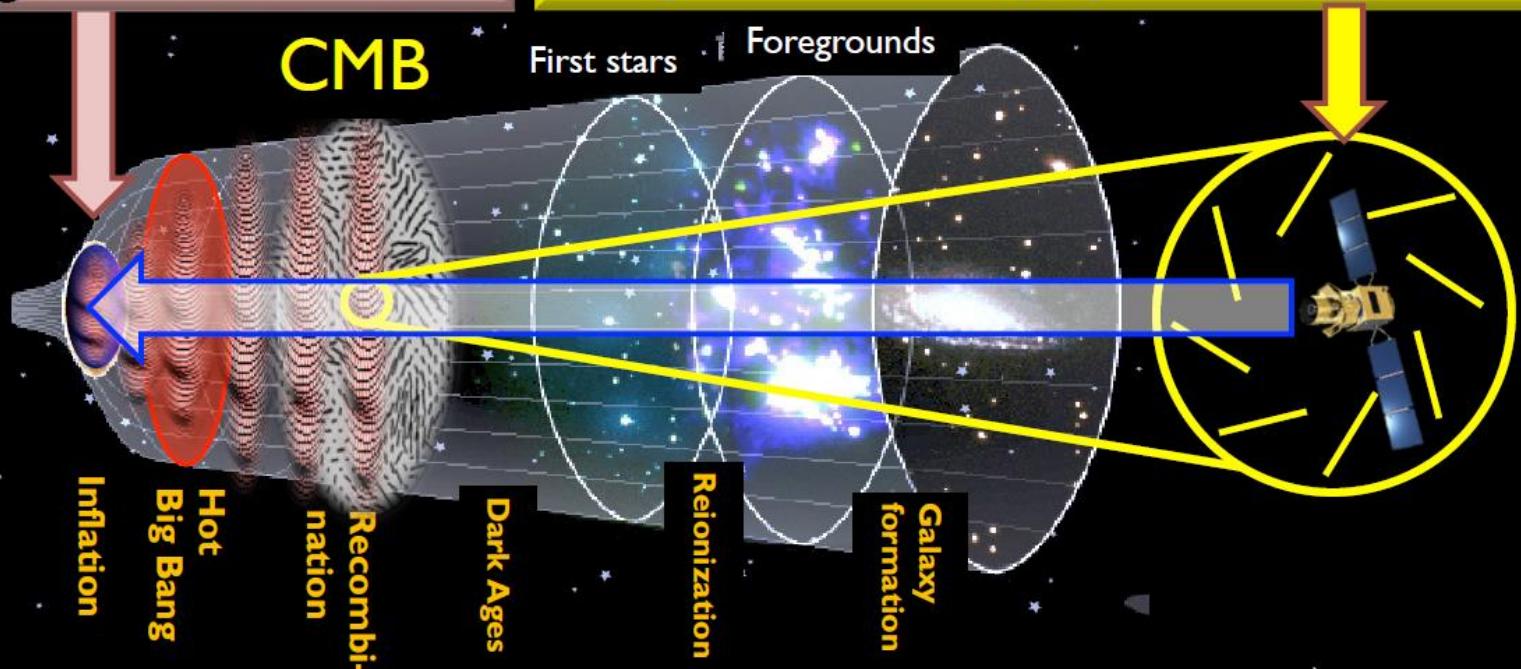
Dark Ages

Reioniza-
tion

Galaxy
formation

Cosmic age 10^{-36} sec? 380000yr

13.8Gyr



JAXA

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T. Matsumura
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R. Keskitalo
T. Kisner
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LiteBIRD working group

X-ray astrophysicists

JAXA engineers

CMB experimenters



IR astronomers

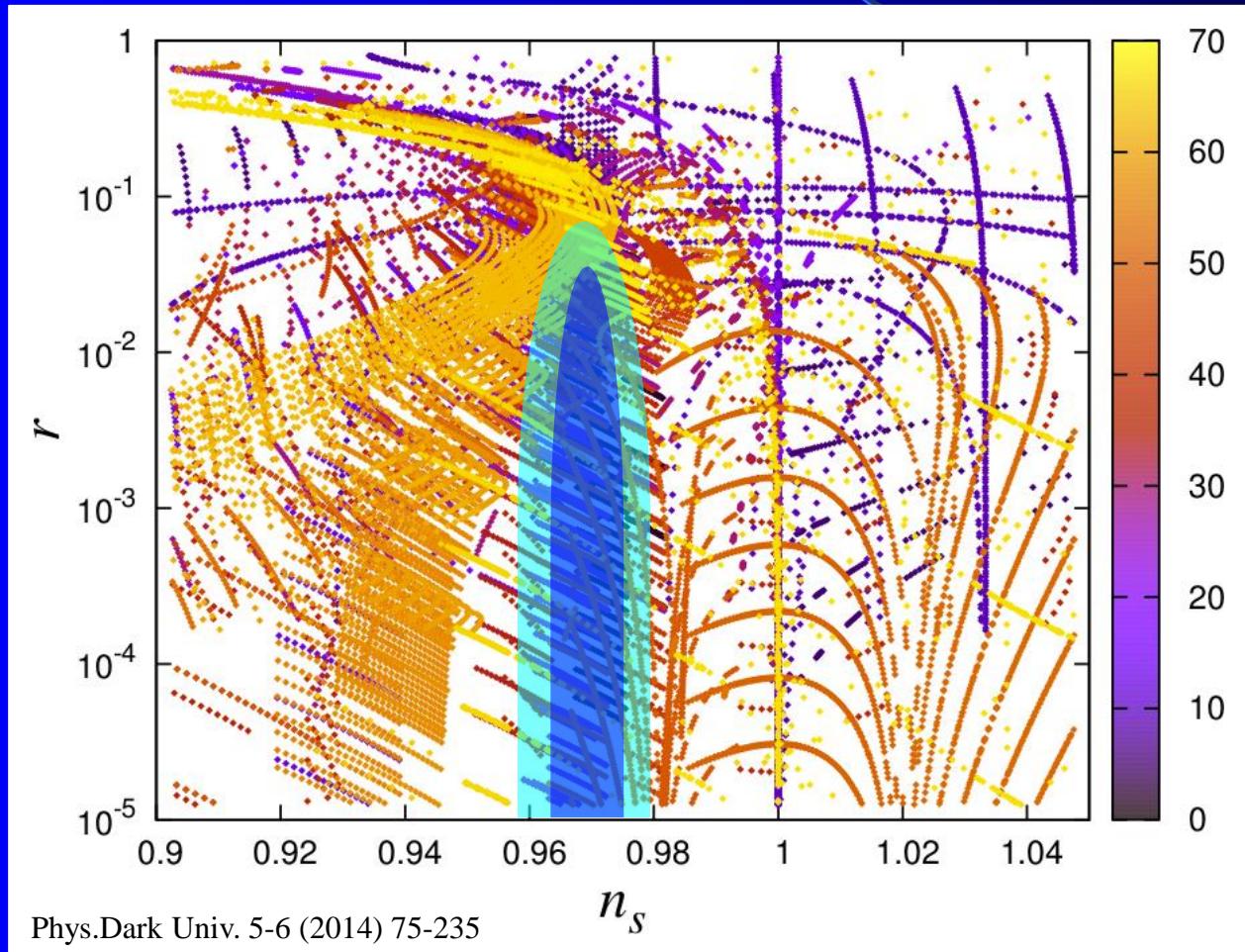
Super-conducting detector developers

UC San Diego

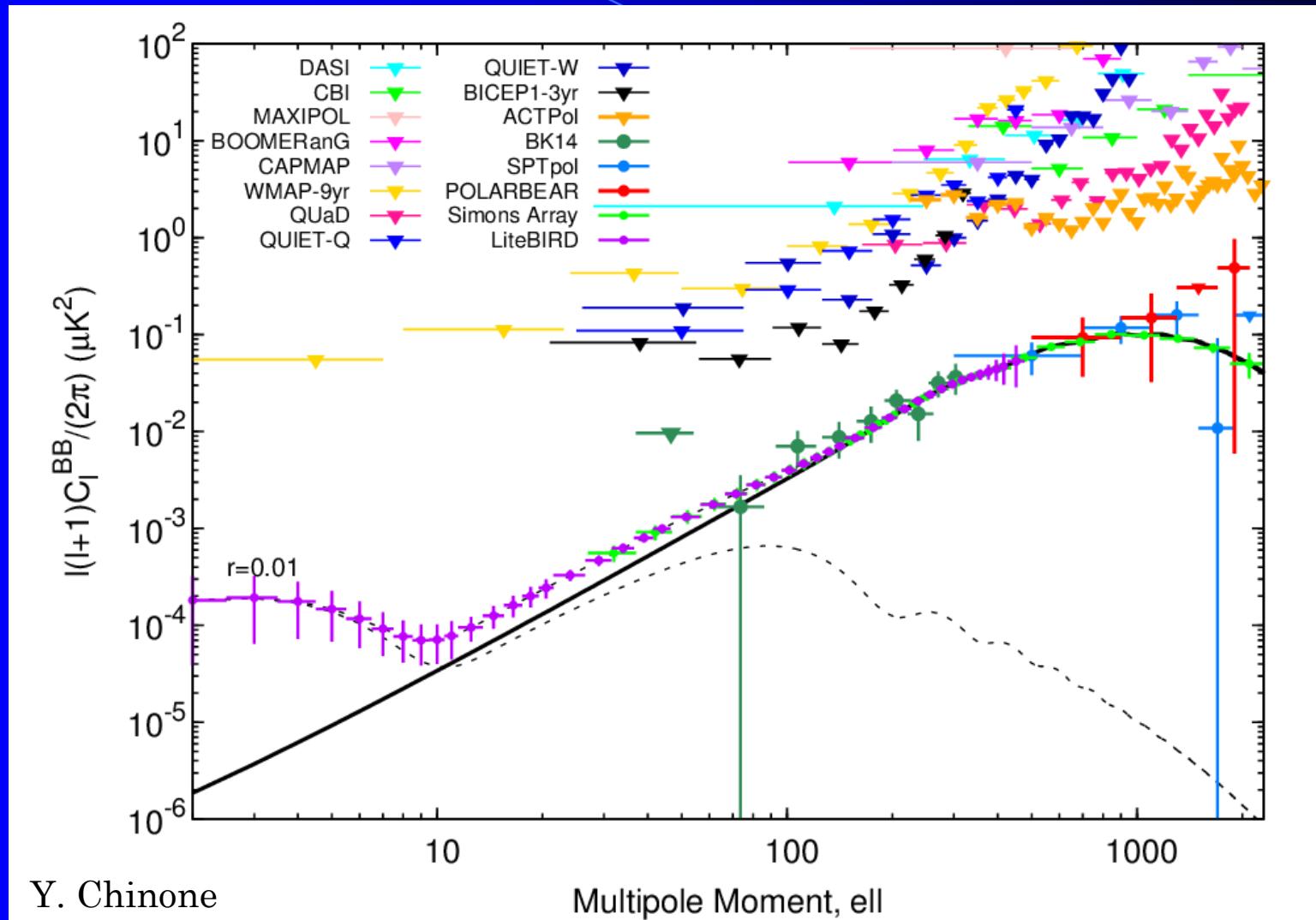
T. Elleot
B. Keating
G. Rebeiz

Inflation models

Single field slow-roll models
with canonical kinetic terms & smooth potentials



Current status of CMB B-mode obs.



Y. Chinone

Multipole Moment, ell

Mission success criterion

$$\sigma_r \leq 0.001$$

- Many models predict $r > 0.01$.
 - 10σ discovery
- Single field models with slow roll condition gives

$$r \simeq 0.002 \left(\frac{60}{N} \right)^2 \left(\frac{\Delta\phi}{m_{pl}} \right)^2$$

N : e-folding, m_{pl} : reduced Planck mass

LiteBIRD error budget

Statistical

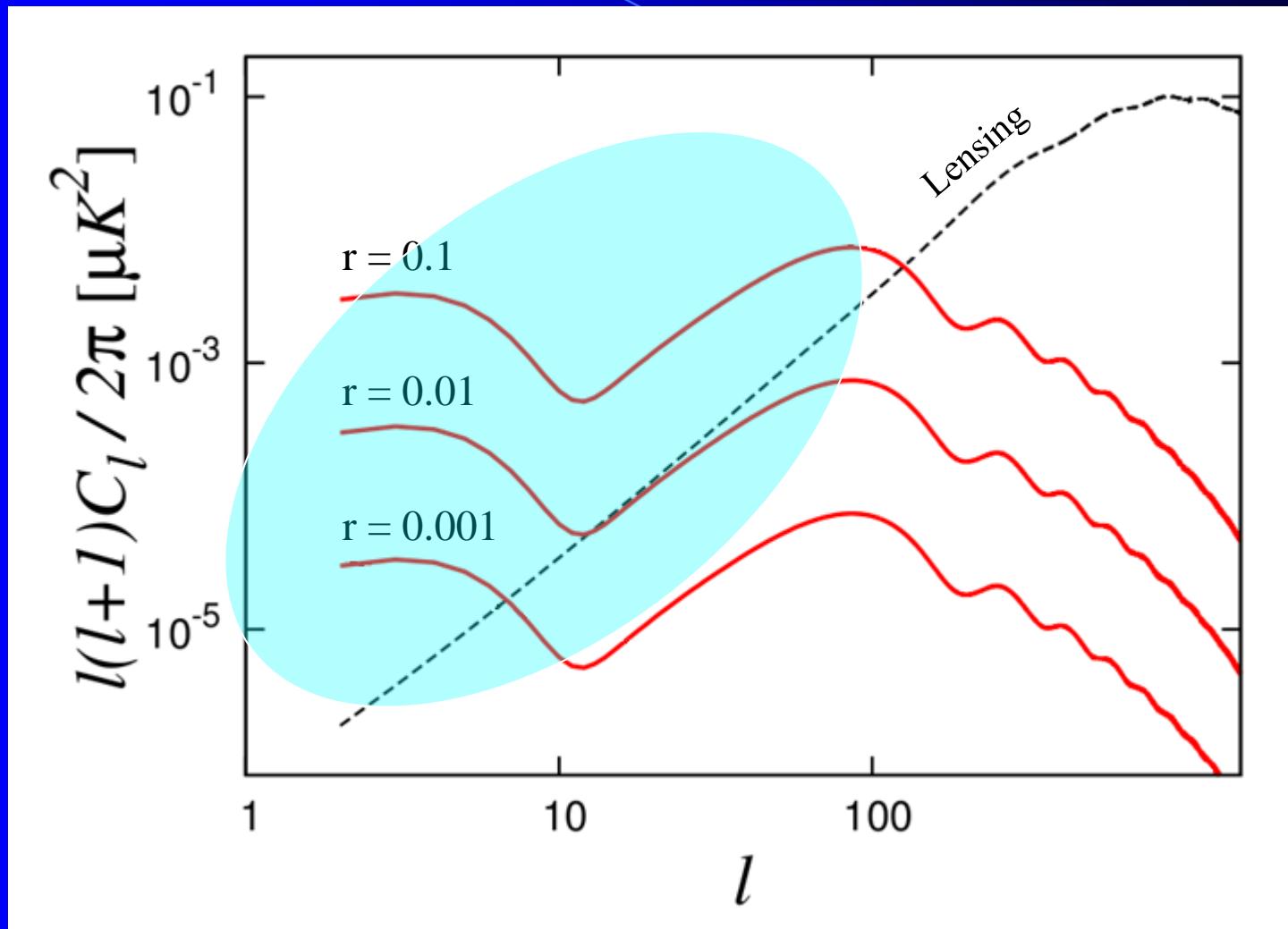
Foreground

Lensing

Systematic

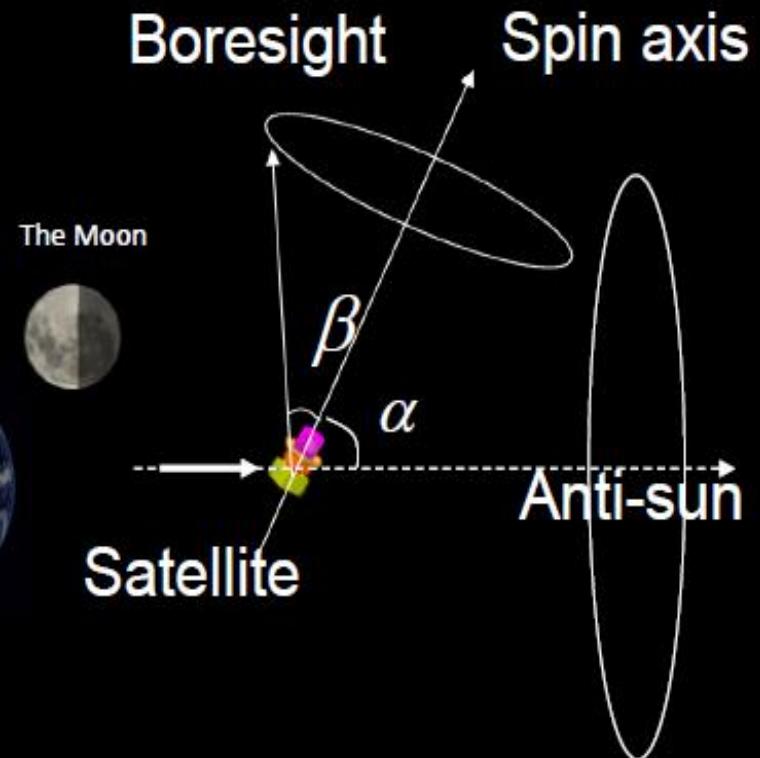
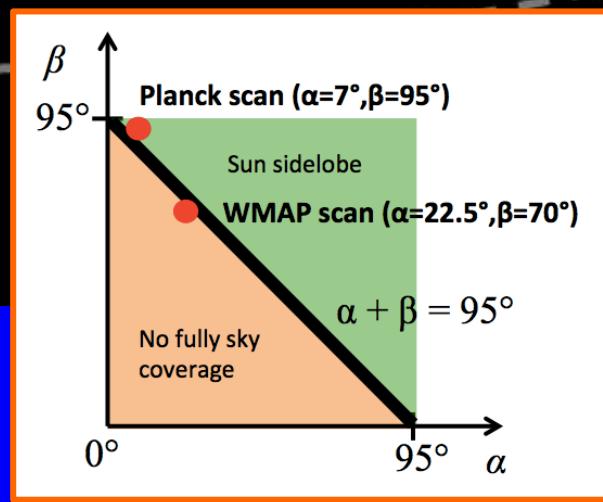
$$\delta r < 0.001$$

Observation strategy – (1)



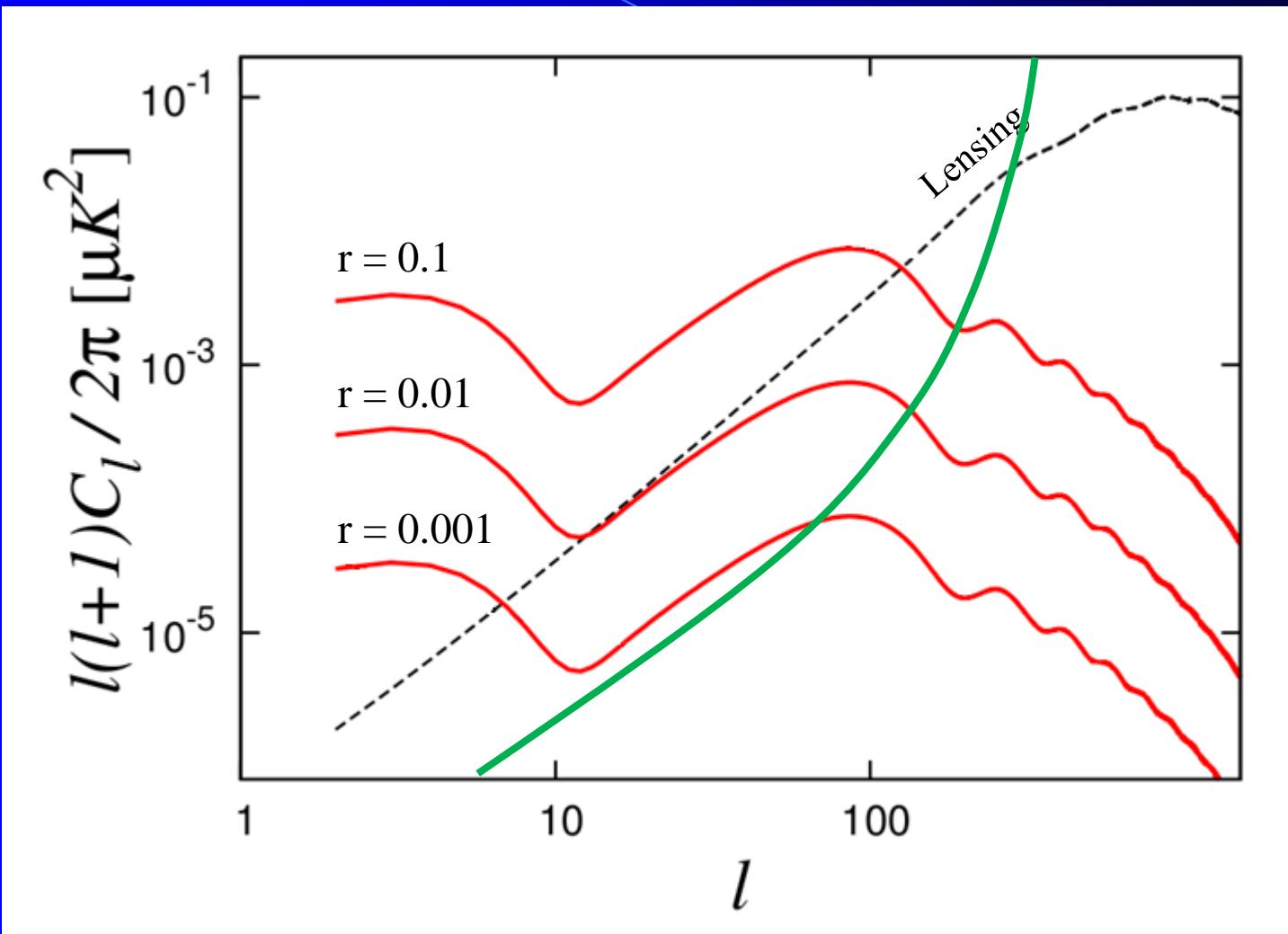
All-sky scan

- Halo orbit at sun-earth L2 point as our baseline
- All-sky scan by combination of precession and spin motion.



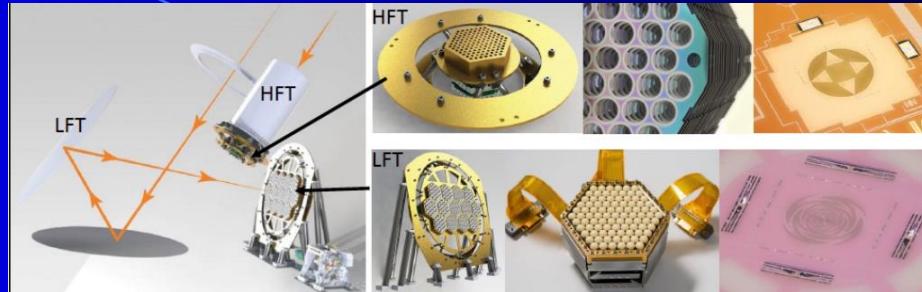
Precession angle $\alpha=65$ deg. 90 min.~1 day/1 rotation
Spin angle $\beta=30$ deg. ~10 min./1 rotation

Observation strategy – (2)



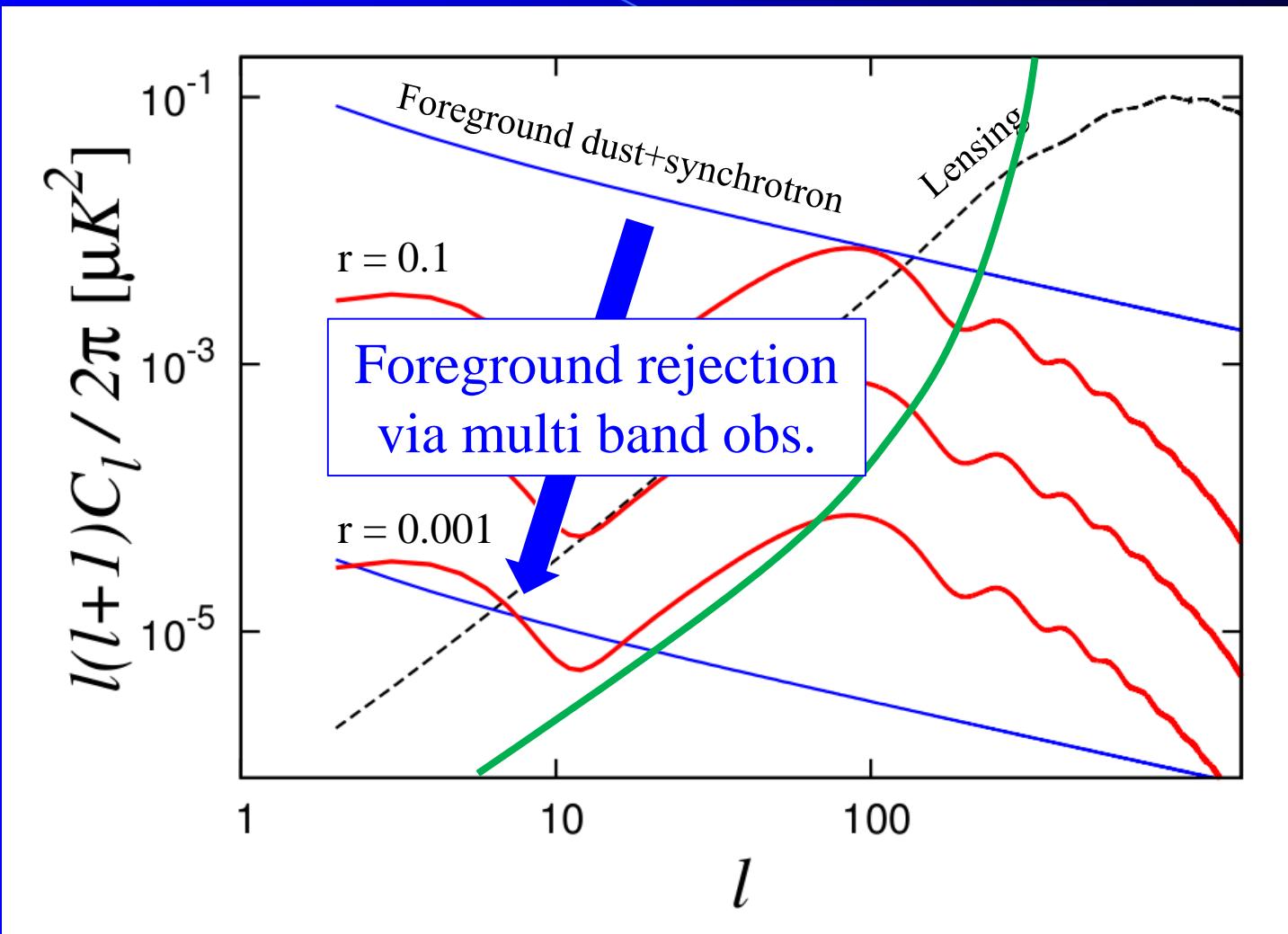
Sensitivity better than lensing floor

Statistical

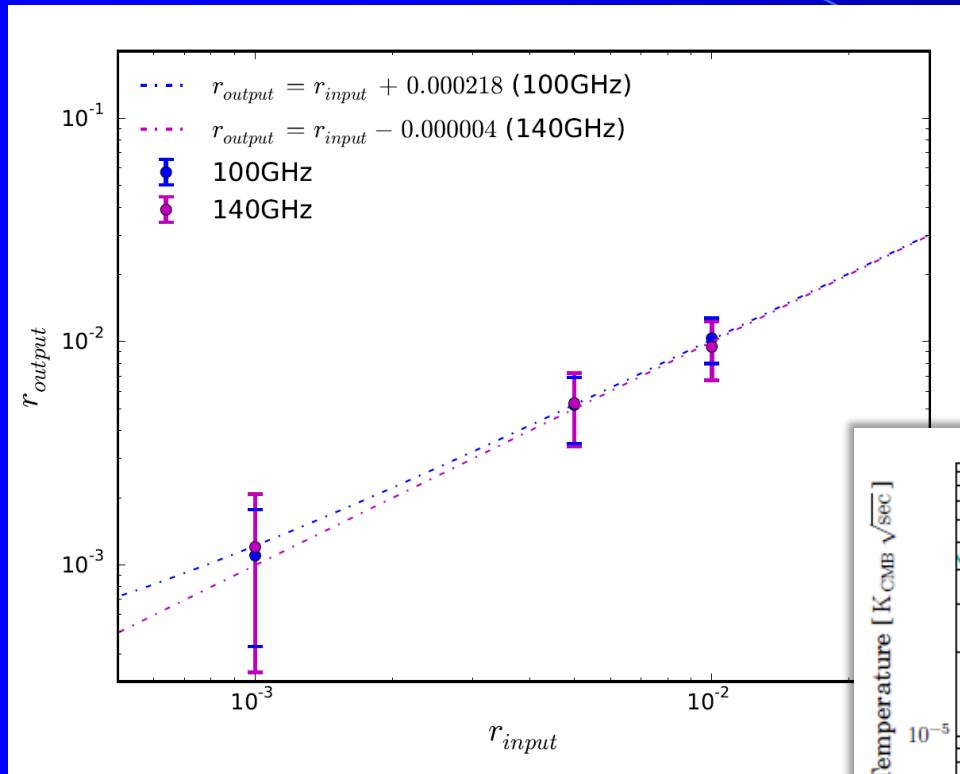


Band (GHz)	Bandwidth ($\Delta\nu/\nu$)	NEP (aW/ $\sqrt{\text{Hz}}$)	NET ($\mu\text{K}\sqrt{\text{s}}$)	N_{bolo}	NET_{arr} ($\mu\text{K}\sqrt{\text{s}}$)	Sensitivity with margin ($\mu\text{K arcmin}$)
40	0.30	7.74	225.9	152	18.3	53.4
50	0.30	7.86	136.9	152	11.1	32.3
60	0.23	7.06	106.2	152	8.6	25.1
68	0.23	7.10	82.9	152	6.7	19.6
78	0.23	7.08	64.7	152	5.2	15.3
89	0.23	7.00	52.4	152	4.3	12.4
100	0.23	8.55	79.7	222	5.3	15.6
119	0.30	9.48	52.5	148	4.3	12.6
140	0.30	8.99	42.3	222	2.8	8.3
166	0.30	8.31	36.2	148	3.0	8.7
195	0.30	7.62	34.1	222	2.3	6.7
235	0.30	6.86	35.8	148	2.9	8.6
280	0.30	9.14	55.4	72	6.5	19.0
338	0.30	8.34	78.0	108	7.5	21.9
402	0.23	6.69	154.4	74	17.9	52.3
Total				2276	3.2	

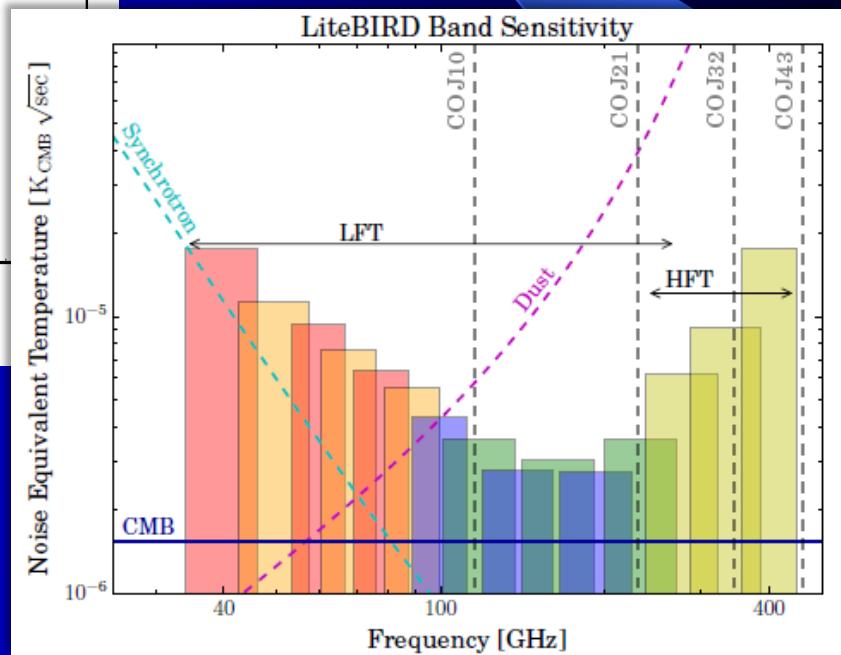
Observation strategy – (3)



Frequency bands from 40GHz to 400GHz

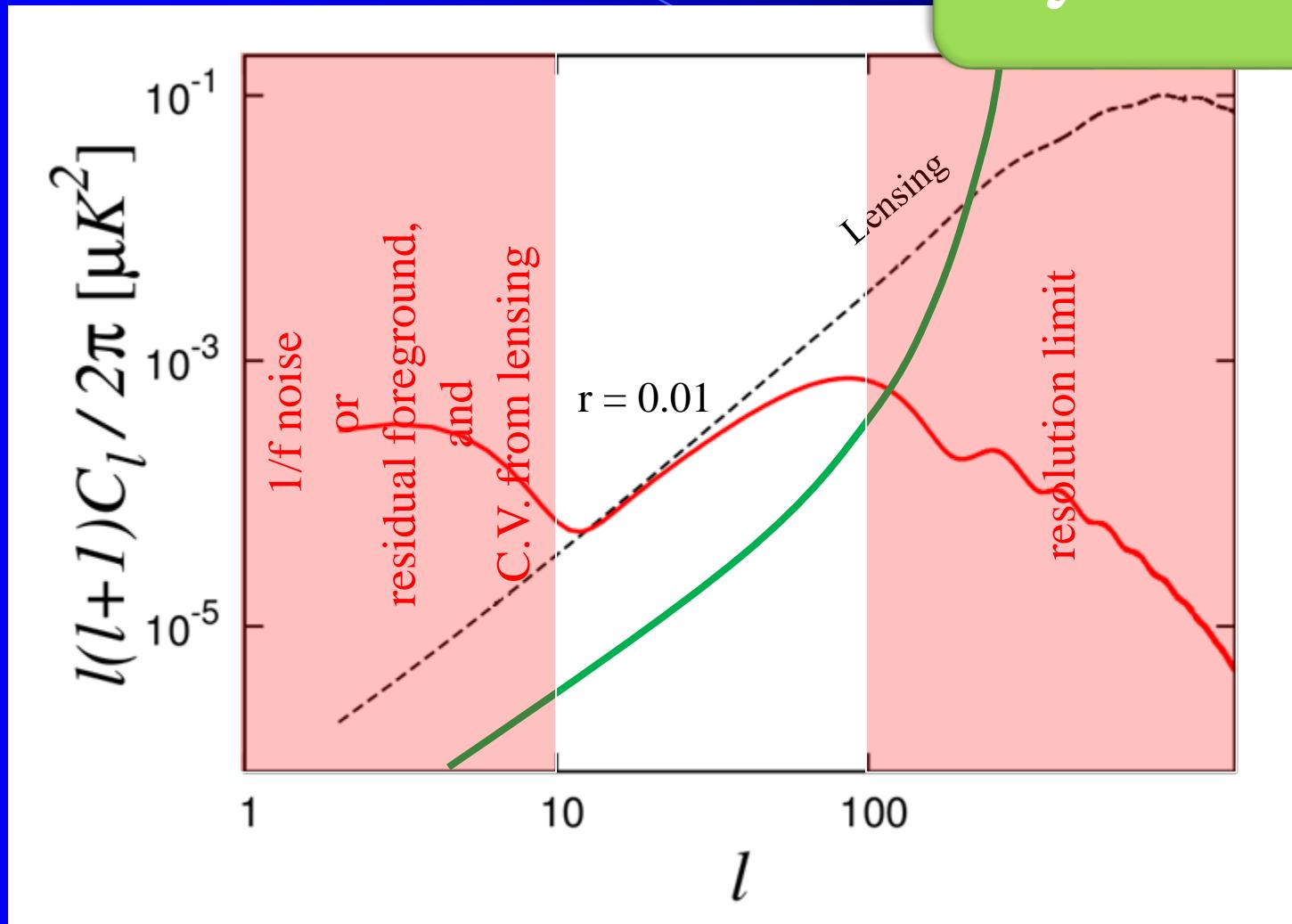


Foreground



Sys. bias, estimated by rule of thumb

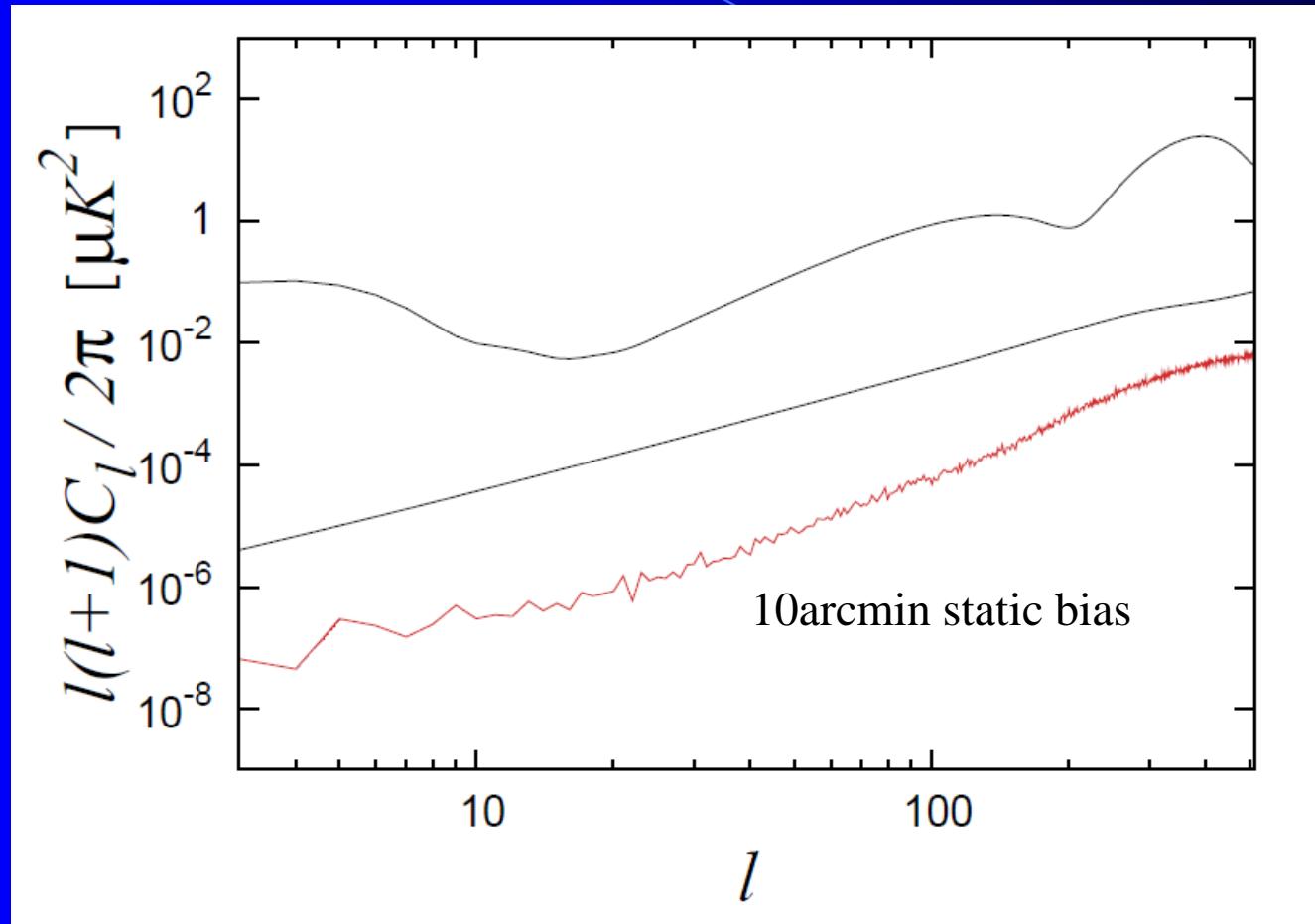
Systematic



Pattern distortion from pointing reconstruction error



Pointing knowledge



Pattern distortion from pol. angle error

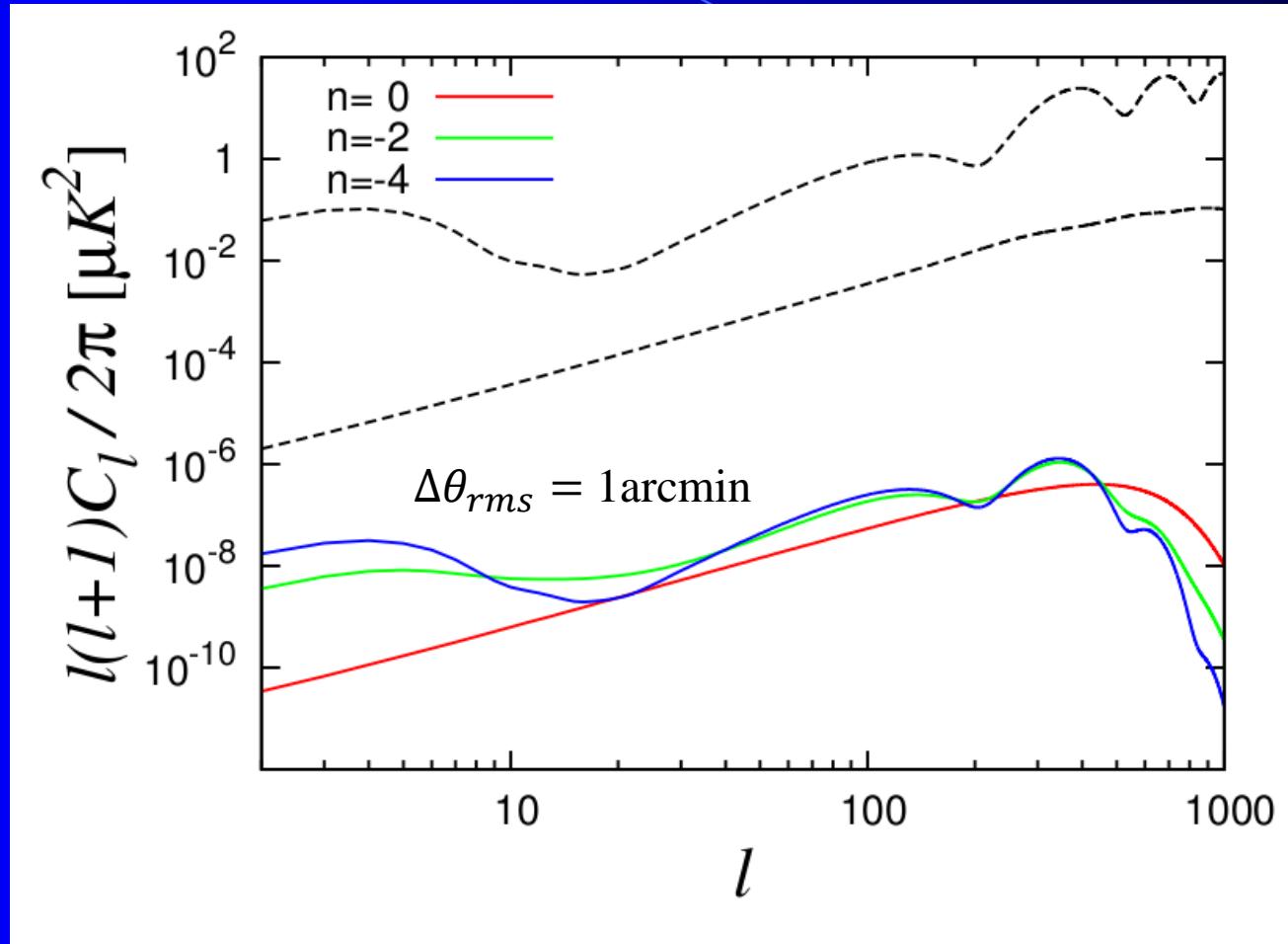
orientation error

attitude, optics, antenna mis-alignment ...

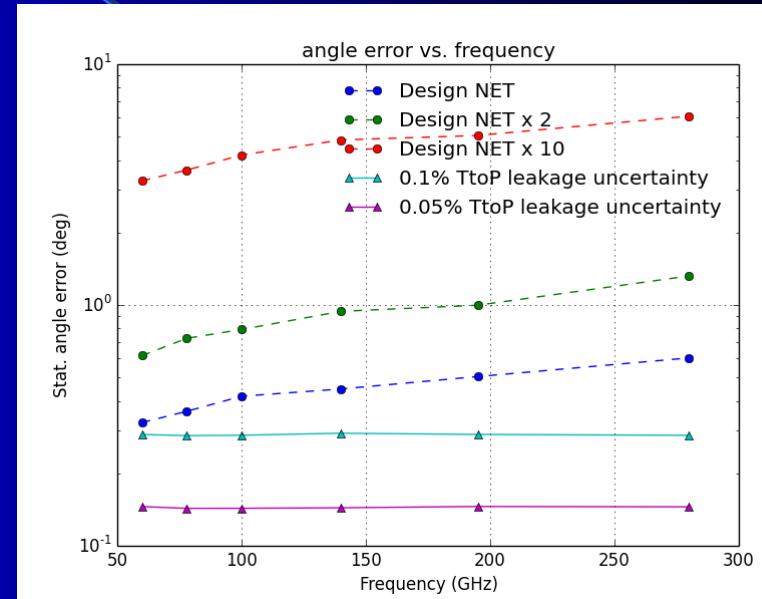
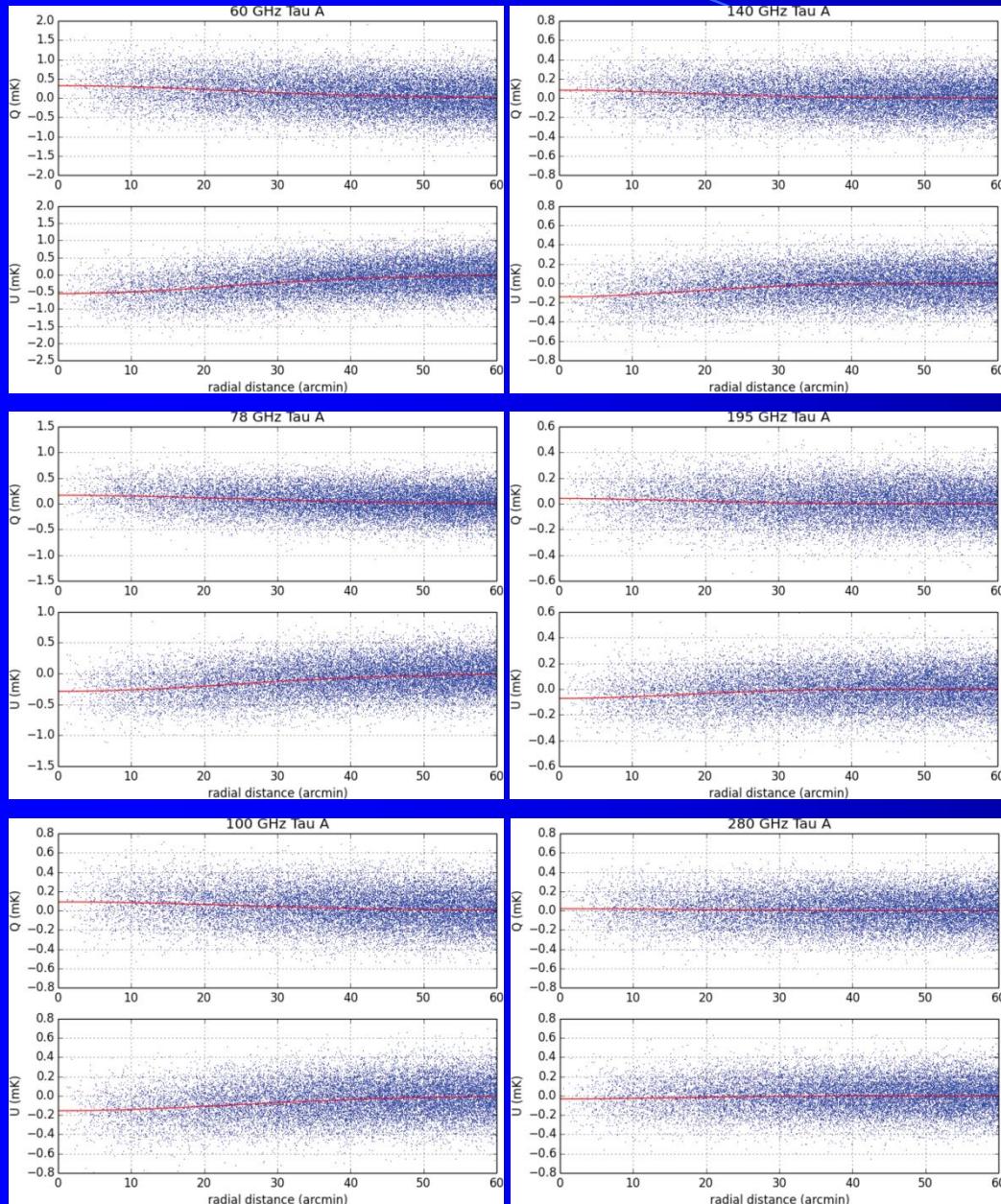
false B-mode

E-mode \Rightarrow B-mode

Polarization angle



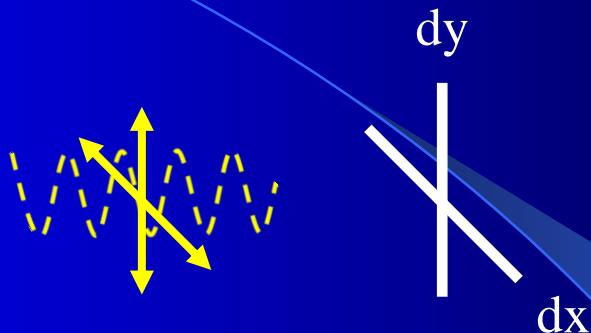
Some speculation about angle calibration



simulation by H.Nishino

Pair differencing pol. measurement

How to measure



$$dx = (\text{non pol.}) + (\text{pol.})$$

$$dy = (\text{non pol.}) - (\text{pol.})$$

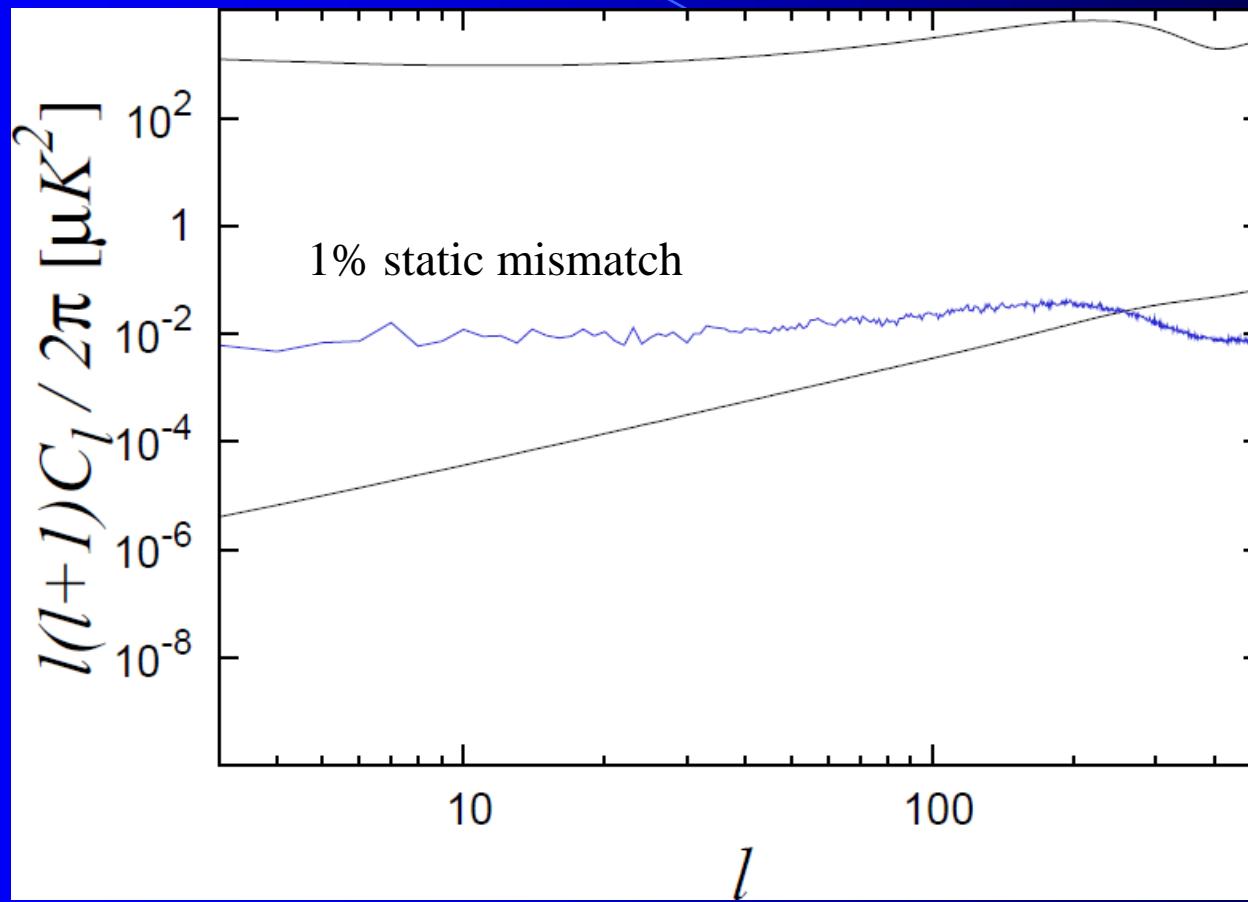
$$\Rightarrow (dx - dy)/2 \quad (\text{pol.})$$

Mismatch gives a residue

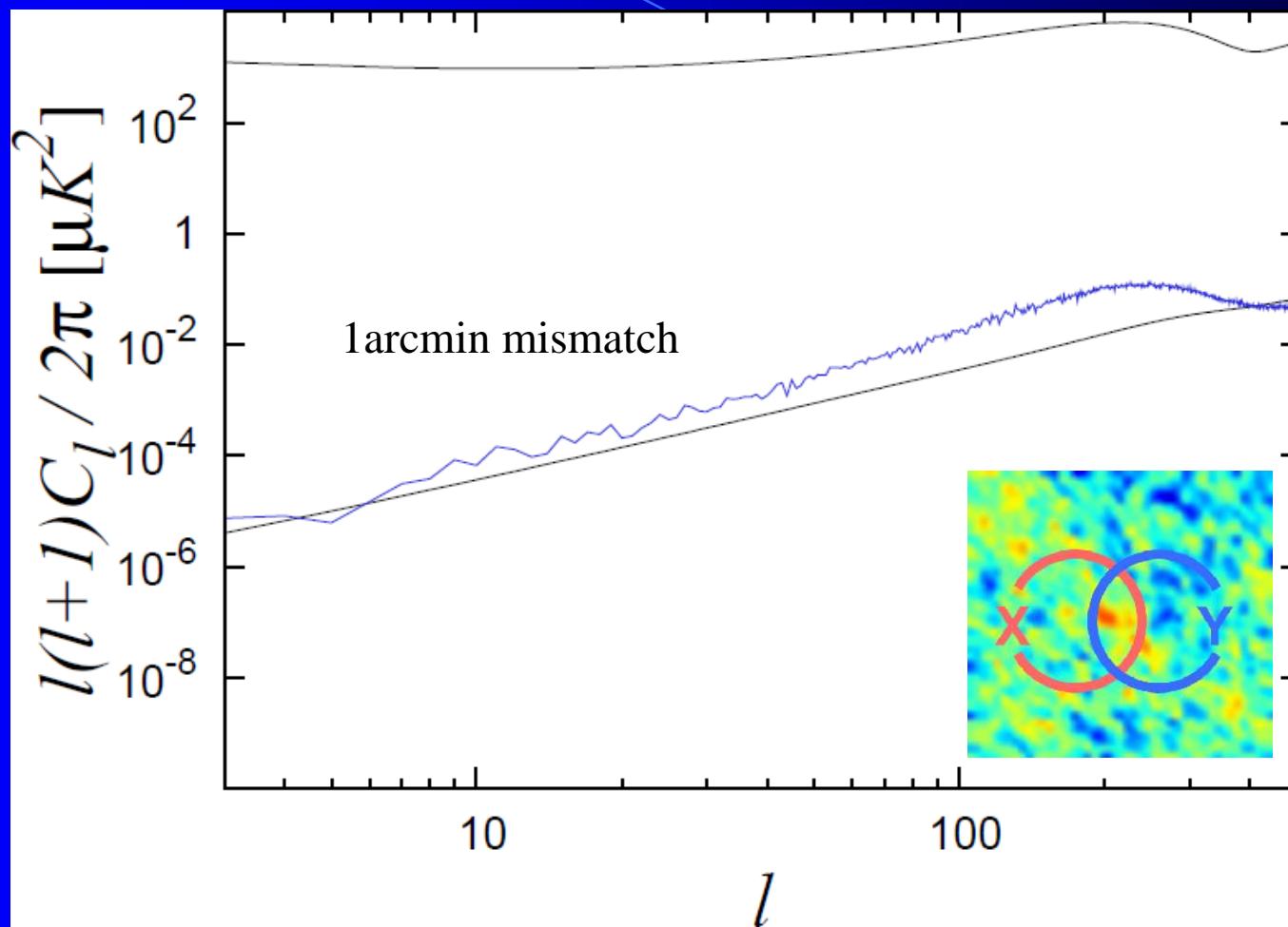
$$(\text{non pol. X}) - (\text{non pol. Y}) \neq 0$$

\Rightarrow false polarization

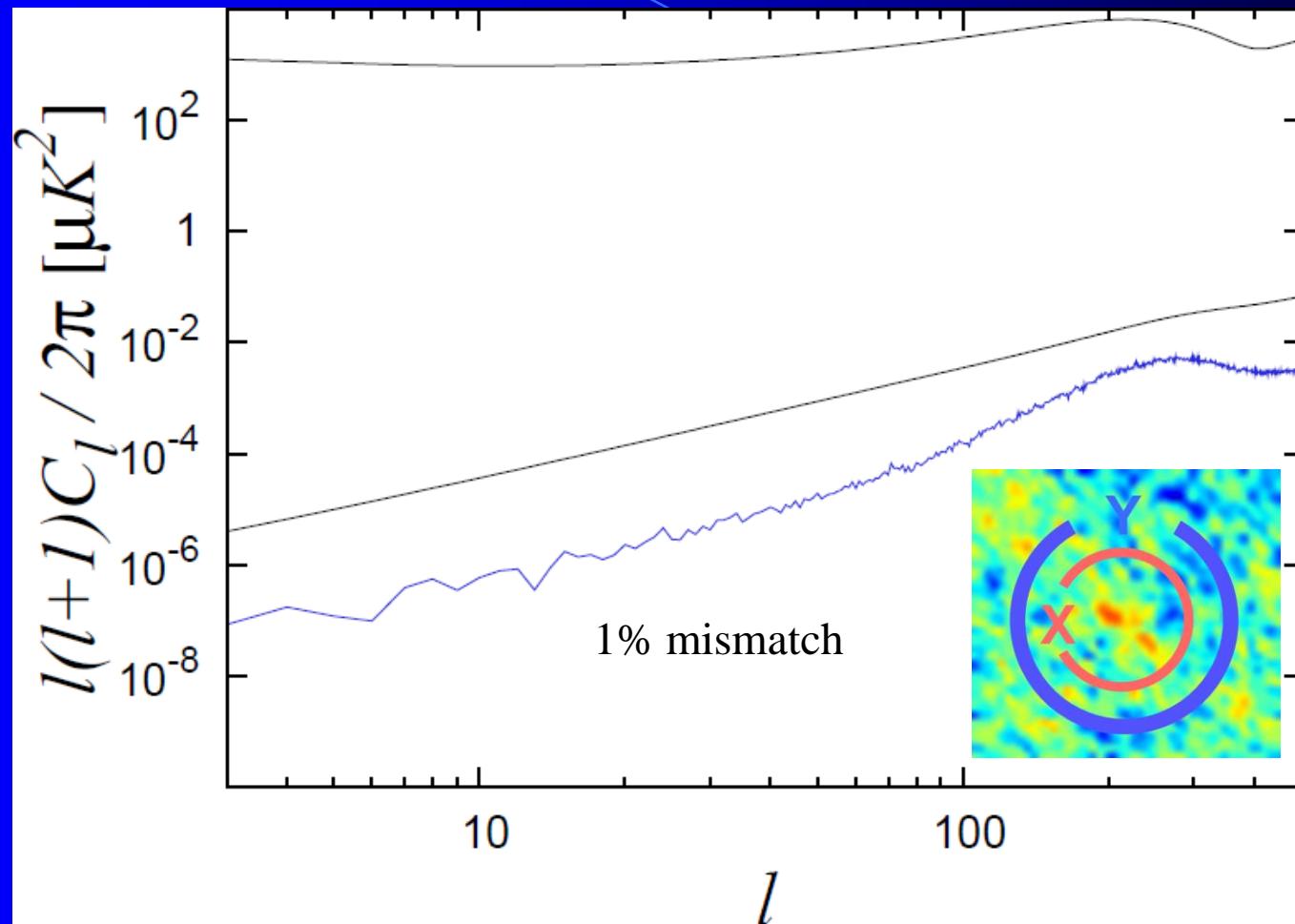
Gain mismatch



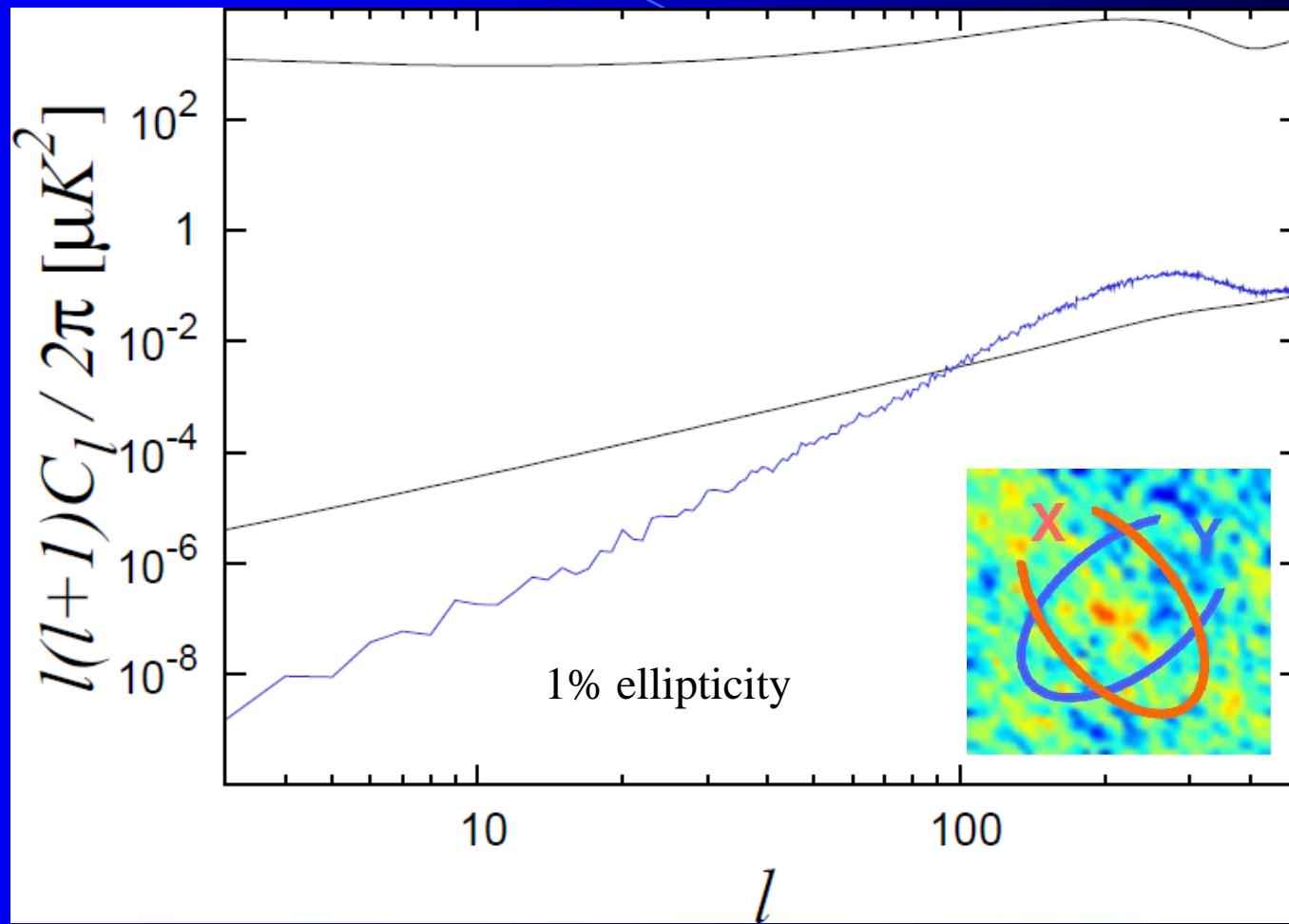
Mainbeam mismatch - beam pointing -



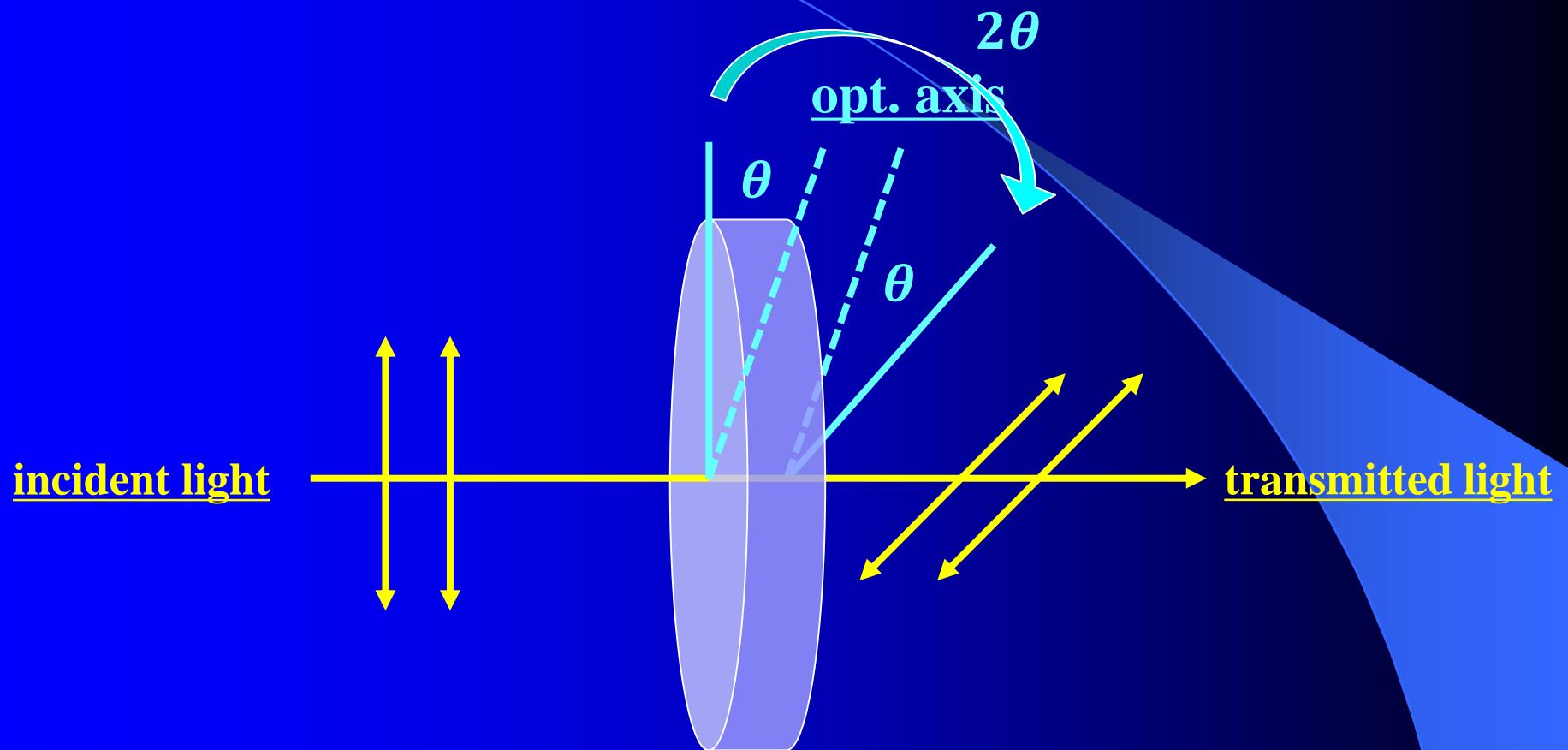
Mainbeam mismatch - beam width -



Mainbeam mismatch - beam ellipticity -

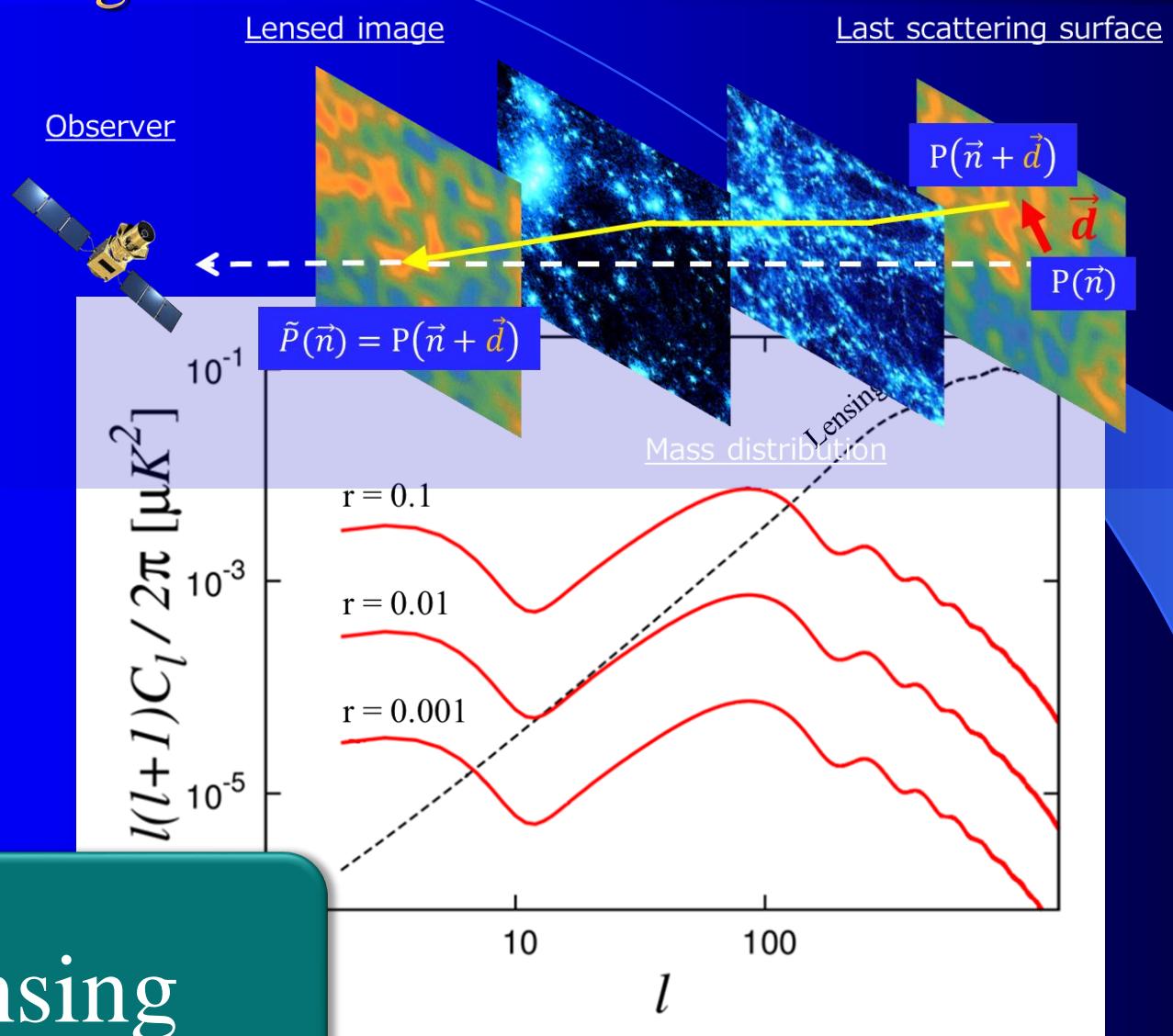


Pol. Modulator (continuously rotating HWP)

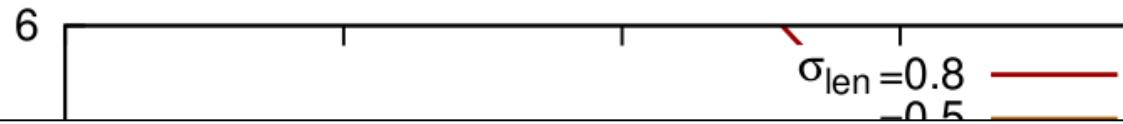


Only pol. component is modulated.

CMB lensing

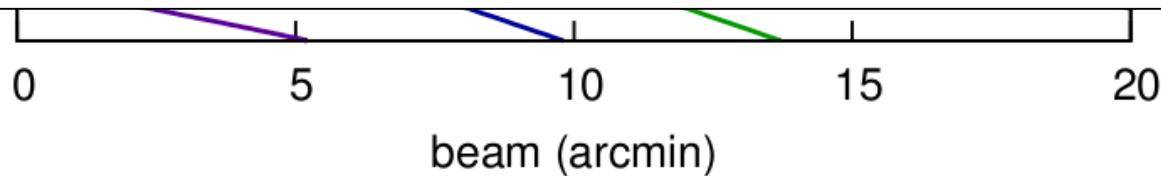


Delensing as extra success

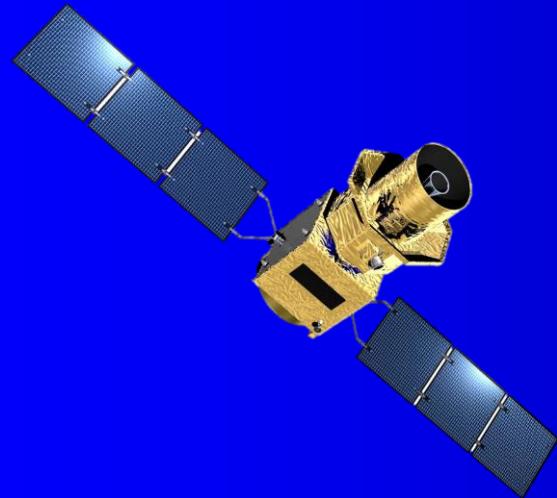


Joint analysis is needed for delensing.

- Large CMB telescope array: e.g. Simons array data
Namikawa and Nagata, JCAP 1409 (2014) 009
- Cosmic infrared background: e.g. Herschel data
Sherwin and Schmittfull, Phys. Rev. D 92, 043005 (2015)
- Radio continuum survey: e.g. SKA data
Namikawa, Yamauchi, Sherwin, Nagata, Phys. Rev. D 93, 043527 (2016)



Synergy with ground-based telescope array



Satellite for ultimate r meas.

$$\sigma(r) < 0.001$$

$$2 \leq l \leq 200$$

Telescope arrays on ground

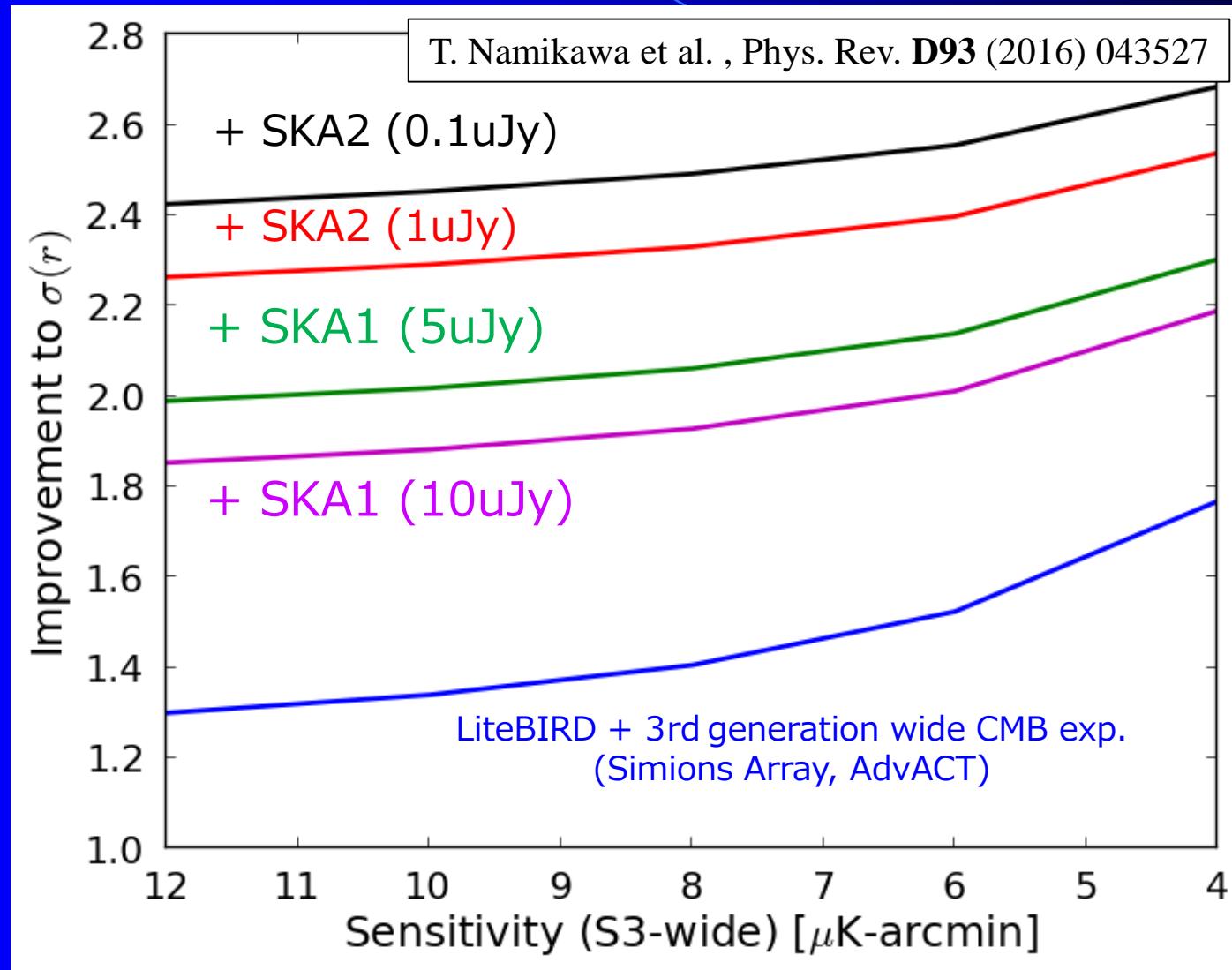
$$30 \leq l \leq \sim 3000$$

e.g. Simons array, CMB-S4

Joint delensing analysis: Namikawa and Nagata, JCAP 1409 (2014) 009

Synergy with SKA

Delensing by use of high-z galaxy map as lensing mass tracer



Summary & current status

- CMB B-mode pol. is the key to the confirmation of cosmic inflation. LiteBIRD is a satellite designed to achieve $\delta r < 0.001$ by precise measurement of B-modes.
- LiteBIRD passed JAXA's initial down-selection and currently in phase-A1.
 - The US LiteBIRD team is working for sub-Kelvin cryogenics and the development of focal plane detectors.
 - The Japan-US JSGs are studying foreground rejection methods by use of multi band obs. and evaluating systematics requirements for designing the satellite and mission instruments.